# Sparse spectral methods for partial differential equations on spherical caps 

Ben Snowball and Sheehan Olver*<br>Department of Mathematics, Imperial College, SE17 3TN, London, UK<br>*Corresponding author: s.olver@imperial.ac.uk

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#### Abstract

In recent years, sparse spectral methods for solving partial differential equations have been derived using hierarchies of classical orthogonal polynomials (OPs) on intervals, disks, disk-slices and triangles. In this work, we extend the methodology to a hierarchy of non-classical multivariate OPs on spherical caps. The entries of discretizations of partial differential operators can be effectively computed using formulae in terms of (non-classical) univariate OPs. We demonstrate the results on partial differential equations involving the spherical Laplacian and biharmonic operators, showing spectral convergence with discretizations that can be made well conditioned using a simple preconditioner.


## 1. Introduction

This paper develops sparse spectral methods for solving linear partial differential equations on certain subsets of the sphere-specifically spherical caps. More precisely, we consider the solution of partial differential equations on the spherical cap $\Omega$ defined by

$$
\Omega:=\left\{(x, y, z) \in \mathbb{R}^{3} \quad \mid \quad \alpha<z<\beta, x^{2}+y^{2}+z^{2}=1\right\},
$$

where $\alpha \in(-1,1)$ and $\beta:=1$.
Remark 1 For simplicity, we focus on the case of a spherical cap, though there is an extension to a spherical band by taking $\beta \in(\alpha, 1)$. The methods presented here translate to the spherical band case by including the necessary adjustments to the weights and recurrence relations we present in this paper. These adjustments make the construction more involved, which is why they are omitted here, but the approach is essentially the same.

We advocate using a basis that is polynomial in Cartesian coordinates, that is, polynomial in $x, y$ and $z$ and orthogonal with respect to a prescribed weight; that is, multivariate orthogonal polynomials (OPs), whose construction was considered in Olver \& Xu (2020). Equivalently, we can think of these as polynomials modulo the vanishing ideal $\left\{x^{2}+y^{2}+z^{2}=1\right\}$, or simply as a linear recombination of spherical harmonics that are orthogonalized on a subset of the sphere. This is in contrast to more standard approaches based on mapping the geometry to a simpler one (e.g., a rectangle or disk) and using OPs in the mapped coordinates (e.g., a basis that is polynomial in the spherical coordinates $\varphi$ and $\theta$ ). The benefit of the new approach is that we do not need to resolve Jacobians, and thereby, we can achieve sparse discretizations for partial differential operators, including those with polynomial variable coefficients. Further, we avoid the singular nature at the poles or as $\alpha$ approaches 0 that such a projection may give, since our new approach yields a smooth polynomial basis for all $\alpha \in[-1,1)$.

There are of course other approaches for solving PDEs on surfaces, such as the closest point method (Macdonald \& Ruuth, 2008, 2010). This involves recasting the PDE as one involving a 'closest point'

[^0]operator in a three-dimensional volume that acts as a 'shell' for the surface. While such an approach is useful for other geometries, it is of low order and does not achieve spectral convergence with sparse discretizations. Further, for our domain of interest, the closest point method does not take advantage of any rotational symmetry and so is unable to achieve optimal complexity with a direct solver.

On the spherical cap, the family of weights we consider are of the form

$$
W^{(a)}(x, y, z):=(z-\alpha)^{a}, \quad \text { for } \quad(x, y, z) \in \Omega,
$$

noting that $W^{(a)}(x, y, z)=0$ for $(x, y, z) \in \partial \Omega$ when $a>0$. The corresponding OPs denoted $Q_{n, k, i}^{(a)}(x, y, z)$, where $n$ denotes the polynomial degree, $0 \leq k \leq n$ and $i \in\{0, \min (1, k)\}$. We define these to be orthogonalized lexicographically, that is,

$$
Q_{n, k, i}^{(a)}(x, y, z)=C_{n, k, i} x^{k-i} y^{i} z^{n-k}+(\text { lower order terms })
$$

where $C_{n, k, i} \neq 0$ and 'lower order terms' includes degree $n$ polynomials of the form $x^{j-i} y^{i} z^{n-j}$ where $j<k$. The precise normalization arises from their definition in terms of one-dimensional OPs in Definition 2.

We consider partial differential operators involving the spherical Laplacian (the Laplace-Beltrami operator): in spherical coordinates

$$
\begin{aligned}
& z=\cos \varphi \\
& x=\sin \varphi \cos \theta=\rho(z) \cos \theta \\
& y=\sin \varphi \sin \theta=\rho(z) \sin \theta
\end{aligned}
$$

where $\rho(z):=\sqrt{1-z^{2}}$, we have

$$
\Delta_{\mathrm{S}}=\frac{1}{\sin \varphi} \frac{\partial}{\partial \varphi}\left(\sin \varphi \frac{\partial}{\partial \varphi}\right)+\frac{1}{\sin ^{2} \varphi} \frac{\partial^{2}}{\partial \theta^{2}}=\frac{1}{\rho} \frac{\partial}{\partial \varphi}\left(\rho \frac{\partial}{\partial \varphi}\right)+\frac{1}{\rho^{2}} \frac{\partial^{2}}{\partial \theta^{2}}
$$

i.e., $\Delta_{\mathrm{S}} f(\mathbf{x})=\Delta f\left(\frac{\mathbf{x}}{\|\mathbf{x}\|}\right)$ for $\mathbf{x}:=(x, y, z) \in \mathbb{R}^{3}$. We do so by considering the component operators $\rho \frac{\partial}{\partial \varphi}$ and $\frac{\partial}{\partial \theta}$ applied to OPs with a specific choices of weight so that their discretization is sparse; see Theorem 3.1. Sparsity comes from expanding the domain and range of an operator using different choices of the parameter $a$, a la the ultraspherical spectral method for intervals (Olver \& Townsend, 2013), triangles (Olver et al., 2019) and disk-slices and trapeziums (Snowball \& Olver, 2020) and the related work on sparse discretizations involving Jacobi polynomials on disks (Vasil et al., 2016) and spheres (Lecoanet et al., 2019; Vasil et al., 2019). As in the disk-slice case in two dimensions (Snowball \& Olver, 2020), we use an integration-by-parts argument to deduce the sparsity structure.

The three-dimensional OPs defined here involve the same non-classical (in fact, semi-classical; Magnus, 1995, Section 5) one-dimensional OPs as those outlined for the disk-slice, and so methods for calculating these one-dimensional OP recurrence coefficients and integrals have already been outlined (Snowball \& Olver, 2020). In particular, by exploiting the connection with these one-dimensional OPs, we can construct discretizations of general partial differential operators of size $(p+1)^{2} \times(p+1)^{2}$ in $O\left(p^{3}\right)$ operations, where $p$ is the total polynomial degree. This clearly compares favourably to proceeding in a naïve approach where one would require $O\left(p^{6}\right)$ operations.

Note that we consider partial differential operators that are not necessarily rotational invariant: for example, one can use these techniques for Schrödinger operators $\Delta_{\mathrm{S}}+v(x, y, z)$ where $v$ is first approximated by a polynomial. A nice feature is that if the partial differential operator is invariant with respect to rotation around the $z$ axis (e.g., a Schrödinger operator with potential $v(z)$ ) the discretization decouples and can be reordered as a block-diagonal matrix. This improves the complexity further to an optimal $O\left(p^{2}\right)$, which is demonstrated in Fig. 5 with $v(x, y, z)=\cos z$.

An overview of the paper is as follows.
Section 2: we present our definition of a (one-parameter) family of three-dimensional OPs on the spherical cap domain $\Omega$, by combining one-dimensional OPs on the interval ( $\alpha, 1$ ) with Chebyshev polynomials, to form three-dimensional OPs on the spherical cap surface. We show that these families will lead to sparse Jacobi operators for multiplication by $x, y, z$ and demonstrate how to obtain the threedimensional OPs.
Section 3: we define several partial differential operators such as spherical Laplacians and show that these will be sparse when applied to a suitable choice of expansions in bases built from OPs on the spherical cap. We can exactly calculate the non-zero entries of these sparse operators using the quadrature rule associated with the non-classical one-dimensional OPs.
Section 4: we derive a quadrature rule on the spherical cap that can be used to expand a function in the OP basis up to a given order $N$ and demonstrate how to evaluate a function using the Clenshaw algorithm using the coefficients of its expansion.
Section 5: we demonstrate the proposed technique for solving differential equations on the spherical cap such as the Poisson equation, variable coefficient Helmholtz equation and Biharmonic equation.

The code that allows one to produce the numerical examples in this paper is publicly available as a Julia package ${ }^{1}$ to partner the ApproxFun package (Olver, 2019); however, this package is purely experimental at this stage.

## 2. OPs on spherical caps

In this section, we outline the construction and some basic properties of $Q_{n, k, i}^{(a)}(x, y, z)$.

### 2.1 Explicit construction

We can construct the three-dimensional OPs on $\Omega$ from one-dimensional OPs on the interval $[\alpha, \beta]$ and from Chebyshev polynomials. We do so in terms of Fourier series, which, following Olver \& Xu (2020), we write here as OPs in $x$ and $y$.
Definition 1 Define the unit circle $\omega:=\left\{\mathbf{x}=(x, y) \in \mathbb{R}^{2} \mid x^{2}+y^{2}=1\right\}$, and define the parameter $\theta$ for each $(x, y) \in \omega$ by $x=\cos \theta, y=\sin \theta$. Define the polynomials $\left\{Y_{k, i}\right\}$ for $k=0,1, \ldots, i=0,1$ on $(x, y) \in \omega$ by

$$
\begin{aligned}
& Y_{0,0}(\mathbf{x}) \equiv Y_{0,0}(x, y):=Y_{0}=: Y_{0,0}(\theta) \\
& Y_{k, 0}(\mathbf{x}) \equiv Y_{k, 0}(x, y):=T_{k}(x)=\cos k \theta=: Y_{k, 0}(\theta), \quad k=1,2,3, \ldots \\
& Y_{k, 1}(\mathbf{x}) \equiv Y_{k, 1}(x, y):=y U_{k-1}(x)=\sin k \theta=: Y_{k, 1}(\theta), \quad k=1,2,3, \ldots,
\end{aligned}
$$

[^1]where $Y_{0}:=\frac{\sqrt{2}}{2}$ and $T_{k}, U_{k-1}$ are the standard Chebyshev polynomials on the interval $[-1,1]$. The $\left\{Y_{k, i}\right\}$ are orthonormal with respect to the inner product
$$
\langle p, q\rangle_{Y}:=\frac{1}{\pi} \int_{0}^{2 \pi} p(\mathbf{x}(\theta)) q(\mathbf{x}(\theta)) \mathrm{d} \theta
$$

Note that we have defined $Y_{0}$ so as to ensure orthonormality.
Proposition 1 (Olver \& Xu, 2020). Let $w:(\alpha, \beta) \rightarrow \mathbb{R}$ be a weight function. For $n=0,1,2, \ldots$, let $\left\{r_{n, k}\right\}$ be polynomials orthogonal with respect to the weight $\rho(x)^{2 k} w(x)$ where $0 \leq k \leq n$. Then the three-dimensional polynomials defined on $\Omega$

$$
Q_{n, k, i}(x, y, z):=r_{n-k, k}(z) \rho(z)^{k} Y_{k, i}\left(\frac{x}{\rho(z)}, \frac{y}{\rho(z)}\right)
$$

for $i \in 0,1,0 \leq k \leq n, n=0,1,2, \ldots$ are OPs with respect to the inner product

$$
\begin{aligned}
\langle p, q\rangle:= & \int_{\Omega} p(x, y, z) q(x, y, z) w(z) \mathrm{d} A \\
= & \int_{0}^{\cos ^{-1}(\alpha)} \int_{0}^{2 \pi} p(\sin \varphi \cos \theta, \sin \varphi \sin \theta, \cos \varphi) q(\sin \varphi \cos \theta, \sin \varphi \sin \theta, \cos \varphi) \times \\
& w(\cos \varphi) \sin \varphi \mathrm{d} \theta \mathrm{~d} \varphi \\
= & \int_{\alpha}^{1} \int_{0}^{2 \pi} p(\rho(z) \cos \theta, \rho(z) \sin \theta, z) q(\rho(z) \cos \theta, \rho(z) \sin \theta, z) w(z) \mathrm{d} \theta \mathrm{~d} z
\end{aligned}
$$

on $\Omega$, where $\mathrm{d} A=\sin \varphi \mathrm{d} \theta \mathrm{d} \varphi$ is the uniform spherical measure on $\Omega$.
OPs are defined as such by the condition that each polynomial is orthogonal to each other, with respect to the given inner product. In particular, they are orthogonal with respect to all lower degree polynomials.

For the spherical cap, we can use Proposition 1 to create our one-parameter family of OPs. We first introduce notation for our family of non-classical univariate OPs that will be used as the $r_{n}$ polynomials above.

Definition 2 (Snowball \& Olver, 2020). Let $w_{R}^{(a, b)}(x)$ be a weight function on the interval $(\alpha, 1)$ given by

$$
w_{R}^{(a, b)}(x):=(x-\alpha)^{a} \rho(x)^{b}
$$

and define the associated inner product by

$$
\begin{equation*}
\langle p, q\rangle_{w_{R}^{(a, b)}}:=\frac{1}{\omega_{R}^{(a, b)}} \int_{\alpha}^{1} p(x) q(x) w_{R}^{(a, b)}(x) \mathrm{d} x \tag{1}
\end{equation*}
$$

where

$$
\begin{equation*}
\omega_{R}^{(a, b)}:=\int_{\alpha}^{1} w_{R}^{(a, b)}(x) \mathrm{d} x \tag{2}
\end{equation*}
$$

is a normalizing constant. Denote the two-parameter family of orthonormal polynomials on $[\alpha, \beta]$ by $\left\{R_{n}^{(a, b)}\right\}$, orthonormal with respect to the inner product defined in equation (1). That is,

$$
\begin{equation*}
\left\langle R_{n}^{(a, b)}, R_{m}^{(a, b)}\right\rangle_{w_{R}^{(a, b)}}=\delta_{n, m} \tag{3}
\end{equation*}
$$

We can now define the three-dimensional OPs for the spherical cap.
Definition 3 Define the one-parameter three-dimensional OPs via

$$
\begin{equation*}
Q_{n, k, i}^{(a)}(x, y, z):=R_{n-k}^{(a, 2 k)}(z) \rho(z)^{k} Y_{k, i}\left(\frac{x}{\rho(z)}, \frac{y}{\rho(z)}\right), \quad(x, y, z) \in \Omega . \tag{4}
\end{equation*}
$$

By construction, $\left\{Q_{n, k, i}^{(a)}\right\}$ are orthogonal with respect to the inner product

$$
\begin{aligned}
\langle p, q\rangle_{Q^{(a)}} & :=\int_{\Omega} p(\mathbf{x}, z) q(\mathbf{x}, z) w_{R}^{(a, 0)}(z) \mathrm{d} A \\
& =\int_{\alpha}^{1} \int_{0}^{2 \pi} p(\rho(z) \cos \theta, \rho(z) \sin \theta, z) q(\rho(z) \cos \theta, \rho(z) \sin \theta, z) \mathrm{d} \theta w_{R}^{(a, 0)}(z) \mathrm{d} z
\end{aligned}
$$

with

$$
\begin{equation*}
\left\|Q_{n, k, i}^{(a)}\right\|_{Q^{(a)}}^{2}:=\left\langle Q_{n, k, i}^{(a)}, Q_{n, k, i}^{(a)}\right\rangle_{Q^{(a)}}=\pi \omega_{R}^{(a, 2 k)} \tag{5}
\end{equation*}
$$

We note that the weight $w_{R}^{(a, b)}(z)$ has been used in the construction of two-dimensional OPs on diskslices and trapeziums (Snowball \& Olver, 2020), where a method for obtaining recurrence coefficients and evaluating integrals was established (the weight is in fact semi-classical and is equivalent to a generalized Jacobi weight; Magnus, 1995, Section 5).

### 2.2 Jacobi matrices

We can express the three-term recurrences associated with $R_{n}^{(a, b)}$ as

$$
\begin{equation*}
x R_{n}^{(a, b)}(x)=\beta_{n}^{(a, b)} R_{n+1}^{(a, b)}(x)+\alpha_{n}^{(a, b)} R_{n}^{(a, b)}(x)+\beta_{n-1}^{(a, b)} R_{n-1}^{(a, b)}(x), \tag{6}
\end{equation*}
$$

where the coefficients are calculatable (Snowball \& Olver, 2020). We can use equation (6) to determine the three-dimensional recurrences for $Q_{n, k, i}^{(a)}(x, y, z)$. Importantly, we can deduce sparsity in the recurrence relationships. We first require the following lemma.

Lemma 1 The following identities hold for $k=2,3, \ldots, j=0,1, \ldots$ and $i, h \in\{0,1\}$ :

1) $\int_{0}^{2 \pi} Y_{0} Y_{j, h}(\theta) \cos \theta \mathrm{d} \theta=Y_{0} \pi \delta_{0, h} \delta_{1, j}$
2) $\int_{0}^{2 \pi} Y_{0} Y_{j, h}(\theta) \sin \theta \mathrm{d} \theta=Y_{0} \pi \delta_{1, h} \delta_{1, j}$
3) $\int_{0}^{2 \pi} Y_{1, i}(\theta) Y_{j, h}(\theta) \cos \theta \mathrm{d} \theta=\pi \delta_{i, h}\left(Y_{0} \delta_{0, j}+\frac{1}{2} \delta_{2, j}\right)$
4) $\int_{0}^{2 \pi} Y_{1, i}(\theta) Y_{j, h}(\theta) \sin \theta \mathrm{d} \theta=\pi \delta_{|i-1|, h}\left((-1)^{i+1} Y_{0} \delta_{0, j}+(-1)^{i} \frac{1}{2} \delta_{2, j}\right)$
5) $\int_{0}^{2 \pi} Y_{k, i}(\theta) Y_{j, h}(\theta) \cos \theta \mathrm{d} \theta=\frac{1}{2} \pi \delta_{i, h}\left(\delta_{k-1, j}+\delta_{k+1, j}\right)$
6) $\int_{0}^{2 \pi} Y_{k, i}(\theta) Y_{j, h}(\theta) \sin \theta \mathrm{d} \theta=\frac{1}{2} \pi \delta_{|i-1|, h}\left((-1)^{i+1} \delta_{k-1, j}+(-1)^{i} \delta_{k+1, j}\right)$.

Proof. Each follows from the definitions of $Y_{k, i}$ and $Y_{0}$, as well as the relationships

$$
\begin{aligned}
2 \cos k \theta \cos \theta & =\cos (k-1) \theta+\cos (k+1) \theta \\
2 \sin k \theta \cos \theta & =\sin (k-1) \theta+\sin (k+1) \theta \\
2 \cos k \theta \sin \theta & =-\sin (k-1) \theta+\sin (k+1) \theta \\
2 \sin k \theta \sin \theta & =\cos (k-1) \theta-\cos (k+1) \theta
\end{aligned}
$$

Lemma 2 Define

$$
\eta_{k}:= \begin{cases}0 & \text { if } k<0  \tag{7}\\ Y_{0} & \text { if } k=0 \\ \frac{1}{2} & \text { otherwise }\end{cases}
$$

$Q_{n, k, i}^{(a)}(x, y, z)$ satisfy the following recurrences:

$$
\begin{aligned}
x Q_{n, k, i}^{(a)}(x, y, z)= & \alpha_{n, k, 1}^{(a)} Q_{n-1, k-1, i}^{(a)}(x, y, z)+\alpha_{n, k, 2}^{(a)} Q_{n-1, k+1, i}^{(a)}(x, y, z) \\
& +\alpha_{n, k, 3}^{(a)} Q_{n, k-1, i}^{(a)}(x, y, z)+\alpha_{n, k, 4}^{(a)} Q_{n, k+1, i}^{(a)}(x, y, z) \\
& +\alpha_{n, k, 5}^{(a)} Q_{n+1, k-1, i}^{(a)}(x, y, z)+\alpha_{n, k, 6}^{(a)} Q_{n+1, k+1, i}^{(a)}(x, y, z), \\
y Q_{n, k, i}^{(a)}(x, y, z)= & \beta_{n, k, i, 1}^{(a)} Q_{n-1, k-1,|i-1|}^{(a)}(x, y, z)+\beta_{n, k, i, 2}^{(a)} Q_{n-1, k+1,|i-1|}^{(a)}(x, y, z) \\
& +\beta_{n, k, i, 3}^{(a)} Q_{n, k-1,|i-1|}^{(a)}(x, y, z)+\beta_{n, k, i, 4}^{(a)} Q_{n, k+1,|i-1|}^{(a)}(x, y, z) \\
& +\beta_{n, k, i, 5}^{(a)} Q_{n+1, k-1,|i-1|}^{(a)}(x, y, z)+\beta_{n, k, i, 6}^{(a)} Q_{n+1, k+1,|i-1|}^{(a)}(x, y, z), \\
z Q_{n, k, i}^{(a)}(x, y, z)= & \gamma_{n, k, 1}^{(a)} Q_{n-1, k, i}^{(a)}(x, y, z)+\gamma_{n, k, 2}^{(a)} Q_{n, k, i}^{(a)}(x, y, z)+\gamma_{n, k, 3}^{(a)} Q_{n+1, k, i}^{(a)}(x, y, z),
\end{aligned}
$$

for $(x, y, z) \in \Omega$, where

$$
\begin{aligned}
& \alpha_{n, k, 1}^{(a)}:=\eta_{k-1}\left\langle R_{n-k}^{(a, 2 k)}, R_{n-k}^{(a, 2(k-1))}\right\rangle_{w_{R}^{(a, 2 k)}}, \\
& \alpha_{n, k, 2}^{(a)}:=\eta_{k}\left\langle R_{n-k}^{(a, 2 k)}, R_{n-k-2}^{(a, 2(k+1))}\right\rangle_{w_{R}^{(a, 2(k+1))}}, \\
& \alpha_{n, k, 3}^{(a)}:=\eta_{k-1}\left\langle R_{n-k}^{(a, 2 k)}, R_{n-k+1}^{(a, 2(k-1))}\right\rangle_{w_{R}^{(a, 2 k)}}, \\
& \alpha_{n, k, 4}^{(a)}:=\eta_{k}\left\langle R_{n-k}^{(a, 2 k)}, R_{n-k-1}^{(a, 2(k+1))}\right\rangle_{w_{R}^{(a, 2(k+1))},}^{(a,} \\
& \alpha_{n, k, 5}^{(a)}:=\eta_{k-1}\left\langle R_{n-k}^{(a, 2 k)}, R_{n-k+2}^{(a, 2(k-1))}\right\rangle_{w_{R}}^{(a, 2 k),} \\
& \alpha_{n, k, 6}^{(a)}:=\eta_{k}\left\langle R_{n-k}^{(a, 2 k)}, R_{n-k}^{(a, 2(k+1))}\right\rangle_{w_{R}^{(a, 2(k+1))}}^{(a,}, \\
& \beta_{n, k, i, j}^{(a)}:= \begin{cases}-\alpha_{n, k, j}^{(a)} & \text { if }(i=0 \text { and } j \text { is odd) or }(i=1 \text { and } j \text { is even }) \\
\alpha_{n, k, j}^{(a)} & \text { otherwise }\end{cases} \\
& \gamma_{n, k, 1}^{(a)}:=\beta_{n-k-1}^{(a, 2 k),} \quad \gamma_{n, k, 2}^{(a)}:=\alpha_{n-k}^{(a, 2 k),} \quad \gamma_{n, k, 3}^{(a)}:=\beta_{n-k}^{(a, 2 k) .}
\end{aligned}
$$

Remark 2 For $z$ multiplication, note that different Fourier modes do not interact. This is because multiplication by $z$ is invariant with respect to rotation around the $z$-axis.

Proof. The three-term recurrence for multiplication by $z$ follows from equation (6). For the recurrence for multiplication by $x$, since $\left\{Q_{m, j, h}^{(a)}\right\}$ for $m=0, \ldots, n+1, j=0, \ldots, m, h=0,1$ is an orthogonal basis for any degree $n+1$ polynomial on $\Omega$, we can expand

$$
x Q_{n, k, i}^{(a)}(x, y, z)=\sum_{m=0}^{n+1} \sum_{j=0}^{m} \sum_{h=0}^{1} c_{m, j} Q_{m, j, h}^{(a)}(x, y, z)
$$

These coefficients are given by

$$
c_{m, j}=\left\langle x Q_{n, k, i}^{(a)}, Q_{m, j, h}^{(a)}\right\rangle_{Q^{(a)}}\left\|Q_{m, j, h}^{(a)}\right\|_{Q^{(a)}}^{-2},
$$

where we show the non-zero coefficients that result are the $\alpha_{n, k, 1}^{(a)}, \ldots, \alpha_{n, k, 6}^{(a)}$ in the lemma. Recall from equation (5) that $\left\|Q_{m, j, h}^{(a)}\right\|_{Q^{(a)}}^{2}=\pi \omega_{R}^{(a, 2 j)}$. Then, for $m=0, \ldots, n+1, j=0, \ldots, m$, using a change of variables $(\cos \theta \sin \varphi, \sin \theta \sin \varphi, \cos \varphi)=(x, y, z)$ :

$$
\begin{aligned}
& \left\langle x Q_{n, k, i}^{(a)}, Q_{m, j, h}^{(a)}\right\rangle_{Q^{(a)}} \\
& =\int_{\Omega} Q_{n, k, i}^{(a)}(\mathbf{x}, z) Q_{m, j, h}^{(a)}(\mathbf{x}, z) x w_{R}^{(a, 0)}(z) \mathrm{d} A \\
& =\left(\int_{\alpha}^{1} R_{n-k}^{(a, 2 k)}(z) R_{m-j}^{(a, 2 j)}(z) \rho(z)^{k+j+1} w_{R}^{(a, 0)}(z) \mathrm{d} z\right) \cdot\left(\int_{0}^{2 \pi} Y_{k, i}(\theta) Y_{j, h}(\theta) \cos \theta \mathrm{d} \theta\right)
\end{aligned}
$$

$$
\begin{aligned}
& =\left(\int_{\alpha}^{1} R_{n-k}^{(a, 2 k)}(z) R_{m-j}^{(a, 2 j)}(z) w_{R}^{(a, k+j+1)}(z) \mathrm{d} z\right) \cdot\left(\int_{0}^{2 \pi} Y_{k, i}(\theta) Y_{j, h}(\theta) \cos \theta \mathrm{d} \theta\right) \\
& =\frac{1}{2} \pi \delta_{i, h}\left(\eta_{k-1} \delta_{k-1, j}+\eta_{k} \delta_{k+1, j}\right) \int_{\alpha}^{1} R_{n-k}^{(a, 2 k)}(z) R_{m-j}^{(a, 2 j)}(z) w_{R}^{(a, k+j+1)}(z) \mathrm{d} z,
\end{aligned}
$$

where $\delta_{k, j}$ is the standard Kronecker delta function, using Lemma 1. Similarly, for the recurrence for multiplication by $y$, we can expand

$$
y Q_{n, k, i}^{(a)}(x, y, z)=\sum_{m=0}^{n+1} \sum_{j=0}^{m} \sum_{h=0}^{1} d_{m, j} Q_{m, j, h}^{(a)}(x, y, z) .
$$

These coefficients are given by

$$
d_{m, j}=\left\langle y Q_{n, k, i}^{(a)}, Q_{m, j, h}^{(a)}\right\rangle_{Q^{(a)}}\left\|Q_{m, j, h}^{(a)}\right\|_{Q^{(a)}}^{-2}
$$

where we show the non-zero coefficients that result are the $\beta_{n, k, 1}^{(a)}, \ldots, \beta_{n, k, 6}^{(a)}$ in the lemma:

$$
\begin{aligned}
& \left\langle y Q_{n, k, i}^{(a)}, Q_{m, j, h}^{(a)}\right\rangle_{Q^{(a)}} \\
& =\int_{\Omega} Q_{n, k, i}^{(a)}(\mathbf{x}, z) Q_{m, j, h}^{(a)}(\mathbf{x}, z) y w_{R}^{(a, 0)}(z) \mathrm{d} A \\
& =\left(\int_{\alpha}^{1} R_{n-k}^{(a, 2 k)}(z) R_{m-j}^{(a, 2 j)}(z) \rho(z)^{k+j+1} w_{R}^{(a, 0)}(z) \mathrm{d} z\right) \cdot\left(\int_{0}^{2 \pi} Y_{k, i}(\theta) Y_{j, h}(\theta) \sin \theta \mathrm{d} \theta\right) \\
& =\left(\int_{\alpha}^{1} R_{n-k}^{(a, 2 k)}(z) R_{m-j}^{(a, 2 j)}(z) w_{R}^{(a, k+j+1)}(z) \mathrm{d} z\right) \cdot\left(\int_{0}^{2 \pi} Y_{k, i}(\theta) Y_{j, h}(\theta) \sin \theta \mathrm{d} \theta\right) \\
& =\frac{1}{2} \pi \delta_{|i-1|, h}\left[(-1)^{i+1} \eta_{k-1} \delta_{k-1, j}+(-1)^{i} \eta_{k} \delta_{k+1, j}\right] \int_{\alpha}^{1} R_{n-k}^{(a, 2 k)}(z) R_{m-j}^{(a, 2 j)}(z) w_{R}^{(a, k+j+1)}(z) \mathrm{d} z,
\end{aligned}
$$

where again $\delta_{k, j}$ is the standard Kronecker delta function, and we have used Lemma 1.
The recurrences in Lemma 2 lead to (block) Jacobi matrices that correspond to multiplication by $x$, $y$ and $z$. In later sections, we will use an ordering of the OPs so that they are grouped by Fourier mode $k$, which is convenient for the application of differential and other operators to the vector of coefficients of a given function's expansion (some operators will exploit this ordering for operators where Fourier modes do not interact and thus will be block-diagonal). Before that, the ordering we will use in the remainder of this section is convenient for establishing Jacobi operators for multiplication by $x, y$ and $z$, and hence building the OPs and importantly obtaining the associated recurrence coefficient matrices necessary for efficient function evaluation using the Clenshaw algorithm. In practice, it is simply a matter of converting coefficients between the two orderings. To this end, we define our OP-building
ordering as follows. For $n=0,1,2, \ldots$,

$$
\mathbb{Q}_{n}^{(a)}:=\left(\begin{array}{c}
Q_{n, 0,0}^{(a)}(x, y, z) \\
Q_{n, 1,0}^{(a)}(x, y, z) \\
Q_{n, 1,1}^{(a)}(x, y, z) \\
\vdots \\
Q_{n, n, 0}^{(a)}(x, y, z) \\
Q_{n, n, 1}^{(a)}(x, y, z)
\end{array}\right) \in \mathbb{R}^{2 n+1}, \quad \mathbb{Q}^{(a)}:=\left(\begin{array}{c}
\mathbb{Q}_{0}^{(a)} \\
\mathbb{Q}_{1}^{(a)} \\
\mathbb{Q}_{2}^{(a)} \\
\vdots
\end{array}\right)
$$

and set $J_{x}^{(a)}, J_{y}^{(a)}, J_{z}^{(a)}$ as the Jacobi matrices corresponding to

$$
\begin{align*}
J_{x}^{(a)} \mathbb{Q}^{(a)}(x, y, z) & =x \mathbb{Q}^{(a)}(x, y, z), \\
J_{y}^{(a)} \mathbb{Q}^{(a)}(x, y, z) & =y \mathbb{Q}^{(a)}(x, y, z),  \tag{8}\\
J_{z}^{(a)} \mathbb{Q}^{(a)}(x, y, z) & =z \mathbb{Q}^{(a)}(x, y, z),
\end{align*}
$$

where

$$
J_{x / y / z}^{(a)}=\left(\begin{array}{cccccc}
B_{x / y / z, 0}^{(a)} & A_{x / y / z, 0}^{(a)} & & & & \\
C_{x / y / z, 1}^{(a)} & B_{x / y / z, 1}^{(a)} & A_{x / y / z, 1}^{(a)} & & & \\
& C_{x / y / z, 2}^{(a)} & B_{x / y / z, 2}^{(a)} & A_{x / y / z, 2}^{(a)} & & \\
& & C_{x / y / z, 3}^{(a)} & \ddots & \ddots & \\
& & & \ddots & \ddots & \ddots
\end{array}\right) .
$$

While $J_{x}^{(a)}, J_{y}^{(a)}, J_{z}^{(a)}$ are not Jacobi matrices in the classical sense, they are block Jacobi matrices. However, it is useful to simply label them as such. In fact, note that they are in fact banded-blockbanded matrices.

Definition 4 A block matrix $A$ with blocks $A_{i, j}$ has block-bandwidths $(L, U)$ if $A_{i, j}=0$ for $-L \leq j-i \leq U$ and sub-block-bandwidths $(\lambda, \mu)$ if all blocks $A_{i, j}$ are banded with bandwidths $(\lambda, \mu)$. A matrix where the block-bandwidths and sub-block-bandwidths are small compared with the dimensions is referred to as a banded-block-banded matrix.

Each of these Jacobi matrices are then block-tridiagonal (block-bandwidths $(1,1)$ ). For $J_{x}^{(a)}$, the sub-blocks have sub-block-bandwidths $(2,2)$ :

$$
A_{x, n}^{(a)}:=\left(\begin{array}{ccccc}
0 & A_{n, 0,6}^{(a)} & 0 & & \\
A_{n, 1,5}^{(a)} & \ddots & \ddots & & \\
& \ddots & \ddots & \ddots & \\
& & A_{n, n, 5}^{(a)} & 0 & A_{n, n, 6}^{(a)}
\end{array}\right) \in \mathbb{R}^{(2 n+1) \times(2 n+3)}, \quad n=0,1,2, \ldots
$$

$$
\begin{aligned}
& B_{x, n}^{(a)}:=\left(\begin{array}{cccc}
0 & A_{n, 0,4}^{(a)} & & \\
A_{n, 1,3}^{(a)} & \ddots & \ddots & \\
& \ddots & \ddots & A_{n, n-1,4}^{(a)} \\
& & A_{n, n, 3}^{(a)} & 0
\end{array}\right) \in \mathbb{R}^{(2 n+1) \times(2 n+1)} \quad n=0,1,2, \ldots \\
& C_{x, n}^{(a)}:=\left(\begin{array}{cccc}
0 & A_{n, 0,2}^{(a)} & & \\
A_{n, 1,1}^{(a)} & \ddots & \ddots & \\
& \ddots & \ddots & A_{n, n-2,2}^{(a)} \\
& & \ddots & 0 \\
& & & A_{n, n, 1}^{(a)}
\end{array}\right) \in \mathbb{R}^{(2 n+1) \times(2 n-1)}, \quad n=1,2, \ldots,
\end{aligned}
$$

where for $k=1, \ldots, N, n=k, \ldots, N$,

$$
\begin{aligned}
& A_{n, k, j}^{(a)}:=\left(\begin{array}{cc}
\alpha_{n, k, j}^{(a)} & 0 \\
0 & \alpha_{n, k, j}^{(a)}
\end{array}\right) \in \mathbb{R}^{2 \times 2},(k \neq 1 \text { for } j \text { odd }) \\
& A_{n, 0, j}^{(a)}:=\binom{\alpha_{n, 0, j}^{(a)}}{0} \in \mathbb{R}^{1 \times 2}, j \text { even } \\
& A_{n, 1, j}^{(a)}:=\binom{\alpha_{n, 1, j}^{(a)}}{0} \in \mathbb{R}^{2 \times 1}, j \text { odd. }
\end{aligned}
$$

For $J_{y}^{(a)}$, the sub-blocks have sub-block-bandwidths $(3,3)$ :

$$
\begin{aligned}
A_{y, n}^{(a)} & :=\left(\begin{array}{ccccc}
0 & B_{n, 0,6}^{(a)} & 0 & & \\
B_{n, 1,5}^{(a)} & \ddots & \ddots & & \\
& \ddots & \ddots & \ddots & \\
& & B_{n, n, 5}^{(a)} & 0 & B_{n, n, 6}^{(a)}
\end{array}\right) \in \mathbb{R}^{(2 n+1) \times(2 n+3)}, \quad n=0,1,2, \ldots \\
B_{y, n}^{(a)} & :=\left(\begin{array}{cccc}
0 & B_{n, 0,4}^{(a)} & & \\
B_{n, 1,3}^{(a)} & \ddots & \ddots & \\
& \ddots & \ddots & B_{n, n-1,4}^{(a)} \\
& & B_{n, n, 3}^{(a)} & 0
\end{array}\right) \in \mathbb{R}^{(2 n+1) \times(2 n+1)} \quad n=0,1,2, \ldots
\end{aligned}
$$

$$
C_{y, n}^{(a)}:=\left(\begin{array}{cccc}
0 & B_{n, 0,2}^{(a)} & & \\
B_{n, 1,1}^{(a)} & \ddots & \ddots & \\
& \ddots & \ddots & B_{n, n-2,2}^{(a)} \\
& & \ddots & 0 \\
& & & B_{n, n, 1}^{(a)}
\end{array}\right) \in \mathbb{R}^{(2 n+1) \times(2 n-1)}, \quad n=1,2, \ldots
$$

where for $k=1, \ldots, N, n=k, \ldots, N$,

$$
\begin{aligned}
B_{n, k, j}^{(a)} & :=\left(\begin{array}{cc}
0 & \beta_{n, k, 0, j}^{(a)} \\
\beta_{n, k, 1, j}^{(a)} & 0
\end{array}\right) \in \mathbb{R}^{2 \times 2},(k \neq 1 \text { for } j \text { odd }) \\
B_{n, 0, j}^{(a)} & :=\left(\begin{array}{ll}
0 & \beta_{n, 0,0, j}^{(a)}
\end{array}\right) \in \mathbb{R}^{1 \times 2}, j \text { even } \\
B_{n, 1, j}^{(a)} & :=\binom{0}{\beta_{n, 1,1, j}^{(a)}} \in \mathbb{R}^{2 \times 1}, j \text { odd. }
\end{aligned}
$$

For $J_{z}^{(a)}$, the sub-blocks are diagonal, i.e., have sub-block-bandwidths $(0,0)$ :

$$
\begin{aligned}
& A_{z, n}^{(a)}:=\left(\begin{array}{ccccc}
\Gamma_{n, 0,3}^{(a)} & 0 & & \\
0 & \ddots & \ddots & \\
& \ddots & \ddots & \ddots & \\
& & 0 & \Gamma_{n, n, 3}^{(a)} & 0
\end{array}\right) \in \mathbb{R}^{(2 n+1) \times(2 n+3)}, \quad n=0,1,2, \ldots \\
& B_{z, n}^{(a)}:=\left(\begin{array}{cccc}
\Gamma_{n, 0,2}^{(a)} & & & \\
& \ddots & & \\
& & \ddots & \\
& & & \Gamma_{n, n, 2}^{(a)}
\end{array}\right) \in \mathbb{R}^{(2 n+1) \times(2 n+1)} \quad n=0,1,2, \ldots \\
& C_{z, n}^{(a)}:=\left(\begin{array}{cccc}
\Gamma_{n, 0,1}^{(a)} & 0 & & \\
0 & \ddots & \ddots & \\
& \ddots & \ddots & 0 \\
& & \ddots & \Gamma_{n, n-1,1}^{(a)} \\
& & & 0
\end{array}\right) \in \mathbb{R}^{(2 n+1) \times(2 n-1)}, \quad n=1,2, \ldots,
\end{aligned}
$$

where for $k=1, \ldots, N, n=k, \ldots, N$,

$$
\begin{align*}
\Gamma_{n, k, j}^{(a)} & :=\left(\begin{array}{cc}
\gamma_{n, k, j} & 0 \\
0 & \gamma_{n, k, j}
\end{array}\right) \in \mathbb{R}^{2 \times 2},  \tag{9}\\
\Gamma_{n, 0, j}^{(a)} & :=\gamma_{n, 0, j}^{(a)} . \tag{10}
\end{align*}
$$

Note that the sparsity of the Jacobi matrices (in particular the sparsity of the sub-blocks) comes from the natural sparsity of the three-term recurrences of the one-dimensional OPs and the circular harmonics, meaning that the sparsity is not limited to the specific spherical cap and would extend to the spherical band.

### 2.3 Building the OPs

A multivariate analogue of Clenshaw's algorithm for recursive evaluation of the OPs has been established for methods on other domains in two dimensions such as the triangle (Olver et al., 2019) and the disk-slice (Snowball \& Olver, 2020), and we proceed similarly here. Combining each system in equation (8) into a block-tridiagonal system, for any $(x, y, z) \in \Omega$, yields

$$
\left(\begin{array}{cccc}
1 & & & \\
B_{0}-G_{0}(x, y, z) & A_{0} & & \\
C_{1} & B_{1}-G_{1}(x, y, z) & A_{1} & \\
& C_{2} & B_{2}-G_{2}(x, y, z) & \ddots \\
& & \ddots & \ddots
\end{array}\right) \mathbb{Q}^{(a)}(x, y, z)=\left(\begin{array}{c}
Q_{0}^{(a)} \\
0 \\
0 \\
0 \\
\vdots
\end{array}\right),
$$

where we note $Q_{0}^{(a)}:=Q_{0,0,0}^{(a)}(x, y, z) \equiv R_{0}^{(a, 0)} Y_{0}$, and for each $n=0,1,2 \ldots$,

$$
\begin{aligned}
& A_{n}:=\left(\begin{array}{l}
A_{x, n}^{(a)} \\
A_{y, n}^{(a)} \\
A_{z, n}^{(a)}
\end{array}\right) \in \mathbb{R}^{3(2 n+1) \times(2 n+3)}, \quad C_{n}:=\left(\begin{array}{l}
C_{x, n}^{(a)} \\
C_{y, n}^{(a)} \\
C_{z, n}^{(a)}
\end{array}\right) \in \mathbb{R}^{3(2 n+1) \times(2 n-1)} \quad(n \neq 0), \\
& B_{n}:=\left(\begin{array}{l}
B_{x, n}^{(a)} \\
B_{y, n}^{(a)} \\
B_{z, n}^{(a)}
\end{array}\right) \in \mathbb{R}^{3(2 n+1) \times(2 n+1)}, \quad G_{n}(x, y, z):=\left(\begin{array}{l}
x I_{2 n+1} \\
y I_{2 n+1} \\
z I_{2 n+1}
\end{array}\right) \in \mathbb{R}^{3(2 n+1) \times(n+1)} .
\end{aligned}
$$

For each $n=0,1,2 \ldots$, let $D_{n}^{\top}$ be any matrix that is a left inverse of $A_{n}$, i.e., such that $D_{n}^{\top} A_{n}=I_{2 n+3}$. Multiplying our system by the preconditioner matrix that is given by the block diagonal matrix of the $D_{n}^{\top} \mathrm{s}$, we obtain a lower triangular system (Dunkl \& Xu, 2014, p. 78), which can be expanded to obtain the recurrence

$$
\left\{\begin{array}{l}
\mathbb{Q}_{-1}^{(a)}(x, y, z):=0 \\
\mathbb{Q}_{0}^{(a)}(x, y, z):=Q_{0}^{(a)} \\
\mathbb{Q}_{n+1}^{(a)}(x, y, z)=-D_{n}^{\top}\left(B_{n}-G_{n}(x, y, z)\right) \mathbb{Q}_{n}^{(a)}(x, y, z)-D_{n}^{\top} C_{n} \mathbb{Q}_{n-1}^{(a)}(x, y, z), \quad n=0,1,2, \ldots
\end{array}\right.
$$

Note that we can define an explicit $D_{n}^{\top}$ as follows:

$$
D_{n}^{\top}:=\left(\begin{array}{ccccccc}
0 & & 0 & & & \left(\Gamma_{n, 0,3}^{(a)}\right)^{-1} & \\
\\
& \ddots & & \ddots & & & \ddots \\
& & 0 & & 0 & & \\
& & & & & & \\
& & & & \left.\Gamma_{n, n, 3}^{(a)}\right)^{-1} \\
& & & & \boldsymbol{\eta}_{1}^{\top} & & \\
\end{array}\right) \in \mathbb{R}^{(2 n+3) \times 3(2 n+1)}
$$

for $n=1,2, \ldots$ where again $\Gamma_{n, k, 3}^{(a)}$ are defined in equations $(9,10)$ for $k=0, \ldots, n$, and where $\boldsymbol{\eta}_{0}, \boldsymbol{\eta}_{1} \in \mathbb{R}^{3(2 n+1)}$ with entries given by

$$
\begin{aligned}
& \left(\boldsymbol{\eta}_{0}\right)_{j}= \begin{cases}\frac{1}{\beta_{n, n, 1,6}^{(a)}} & j=2(2 n+1) \\
\frac{\beta_{n, n, 1,5}^{(a)}}{\beta_{n, n, 1,6}^{(a)} \gamma_{n, n-1,3}^{(a)}} & j=3(2 n+1)-3 \\
0 & o / w\end{cases} \\
& \left(\boldsymbol{\eta}_{1}\right)_{j}= \begin{cases}\frac{1}{\alpha_{n, n, 6}^{(a)}} & j=2 n+1 \\
\frac{-\alpha_{n, n, 5}^{(a)}}{\alpha_{n, n, 6}^{(a)} \gamma_{n, n-1,3}^{(a)}} & j=3(2 n+1)-2 \text { and } n>1 \\
0 & o / w\end{cases}
\end{aligned}
$$

For $n=0$, we can simply take

$$
D_{0}^{\top}:=\left(\begin{array}{ccc}
0 & 0 & \frac{1}{\gamma_{0,0,3}^{(a)}} \\
\frac{1}{\alpha_{0,0,6}^{(a)}} & 0 & 0 \\
0 & \frac{1}{\beta_{0,0,6}^{(a)}} & 0
\end{array}\right) \in \mathbb{R}^{3 \times 3}
$$

It follows that we can apply $D_{n}^{\top}$ in $O(n)$ complexity and thereby calculate $\mathbb{Q}_{0}^{(a)}(x, y, z)$ through $\mathbb{Q}_{n}^{(a)}(x, y, z)$ in optimal $O\left(n^{2}\right)$ complexity.
Definition 5 The recurrence coefficient matrices associated with the OPs $\left\{Q_{n, k, i}^{(a)}\right\}$ are given by the matrices $A_{n}, B_{n}, C_{n}, D_{n}^{\top}$ for $n=0,1,2, \ldots$ defined above.

## 3. Sparse partial differential operators

In this section, we will derive the entries of spherical partial differential operators applied to our basis, demonstrating their sparsity in the process. To this end, as alluded to in Section 2.2, we introduce new notation for a different ordering of the OP vector, in order to exploit the orthogonality the polynomials
$Y_{k, i}$ will bring and thus ensure the operators will be block-diagonal. Let $N \in \mathbb{N}$, and define

$$
\begin{align*}
& \tilde{\mathbb{Q}}_{N, k}^{(a)}:=\left(\begin{array}{c}
Q_{k, k, 0}^{(a)}(x, y, z) \\
Q_{k, k, 1}^{(a)}(x, y, z) \\
\vdots \\
Q_{N, k, 0}^{(a)}(x, y, z) \\
Q_{N, k, 1}^{(a)}(x, y, z)
\end{array}\right) \in \mathbb{R}^{2(N-k+1)}, \quad k=1, \ldots, N,  \tag{11}\\
& \tilde{\mathbb{Q}}_{N, 0}^{(a)}:=\left(\begin{array}{c}
Q_{0,0,0}^{(a)}(x, y, z) \\
\vdots \\
Q_{N, 0,0}^{(a)}(x, y, z)
\end{array}\right) \in \mathbb{R}^{N+1} .  \tag{12}\\
& \tilde{\mathbb{Q}}_{N}^{(a)}:=\left(\begin{array}{c}
\tilde{\mathbb{Q}}_{N, 0}^{a( } \\
\vdots \\
\tilde{\mathbb{Q}}_{N, N}^{(a)}
\end{array}\right) \in \mathbb{R}^{(N+1)^{2}} \tag{13}
\end{align*}
$$

We further denote the weighted set of OPs on $\Omega$ by

$$
\mathbb{W}_{N}^{(a)}(x, y, z):=w_{R}^{(a, 0)}(z) \tilde{\mathbb{Q}}_{N}^{(a)}(x, y, z)
$$

The operator matrices we derive here act on coefficient vectors that represent a function $f(x, y, z)$ defined on $\Omega$ in spectral space-such a function is approximated by its expansion up to degree $N$ :

$$
f(x, y, z)=\tilde{\mathbb{Q}}_{N}^{(a)}(x, y, z)^{\top} \boldsymbol{f}=\sum_{n=0}^{N} \sum_{k=0}^{n} \sum_{i=0}^{1} f_{n, k, i} Q_{n, k, i}^{(a)}(x, y, z),
$$

where $\boldsymbol{f}=\left(f_{n, k, i}\right)$ is the coefficients vector for the function $f$.
Definition 6 Let $a$ be a nonnegative parameter and $\tilde{a} \geq 2$ be an integer. Define the operator matrices $D_{\varphi}^{(a)}, W_{\varphi}^{(a)}, D_{\theta}, \mathcal{L}^{(a) \rightarrow(a+\tilde{a})}, \mathcal{L}_{W}^{(a) \rightarrow(a-\tilde{a})}, \Delta_{W}^{(1)}$ according to

$$
\begin{aligned}
\rho \frac{\partial f}{\partial \varphi}(x, y, z) & =\widetilde{\mathbb{Q}}_{N}^{(a+1)}(x, y, z)^{\top} D_{\varphi}^{(a)} \mathbf{f}, \\
\rho \frac{\partial}{\partial \varphi}\left[w_{R}^{(a, 0)}(z) f(x, y, z)\right] & =\mathbb{W}_{N}^{(a-1)}(x, y)^{\top} W_{\varphi}^{(a)} \mathbf{f}, \\
\frac{\partial f}{\partial \theta}(x, y, z) & =\widetilde{\mathbb{Q}}_{N}^{(a)}(x, y, z)^{\top} D_{\theta} \mathbf{f}, \\
\Delta_{\mathrm{S}} f(x, y, z) & =\widetilde{\mathbb{Q}}_{N}^{(a+\tilde{a})}(x, y, z)^{\top} \mathcal{L}^{(a) \rightarrow(a+\tilde{a})} \mathbf{f}, \\
\Delta_{\mathrm{S}}\left(w_{R}^{(a, 0)}(z) f(x, y, z)\right) & =\mathbb{W}_{N}^{(a-\tilde{a})}(x, y, z)^{\top} \mathcal{L}_{W}^{(a) \rightarrow(a-\tilde{a})} \mathbf{f}, \quad \text { (for } a \geq 2 \text { only) } \\
\Delta_{\mathrm{S}}\left(w_{R}^{(1,0)}(z) f(x, y, z)\right) & =\widetilde{\mathbb{Q}}_{N}^{(1)}(x, y, z)^{\top} \Delta_{W}^{(1)} \mathbf{f}, \quad \text { (for } a=1 \text { only). }
\end{aligned}
$$

The incrementing and decrementing of parameters as seen here is analogous to other well-known OP families' derivatives, for example the Jacobi polynomials on the interval, as seen in the DLMF (Olver et al., 2010, (18.9.3)), on the triangle (Olver et al., 2020) and on the disk-slice (Snowball \& Olver, 2020). The operators we define here are for partial derivatives with respect to the spherical coordinates $(\varphi, \theta)$, so that we can more easily apply the operators to PDEs on the surface of a sphere (for example, surface Laplacian operator in the Poisson equation). With the OP ordering by Fourier mode $k$ defined in equations $(11,12,13)$ these rotationally invariant operators are block-diagonal, meaning simple and parallelizable practical application.
Theorem 3.1 The operator matrices $D_{\varphi}^{(a)}, W_{\varphi}^{(a)}, D_{\theta}, \mathcal{L}^{(a) \rightarrow(a+\tilde{a})}, \mathcal{L}_{W}^{(a) \rightarrow(a-\tilde{a})}, \Delta_{W}^{(1)}$ from Definition 6 are sparse, with banded-block-banded structure. More specifically,

- $D_{\varphi}^{(a)}$ is block-diagonal with sub-block-bandwidths $(2,4)$;
- $W_{\varphi}^{(a)}$ is block-diagonal with sub-block-bandwidths $(4,2)$;
- $D_{\theta}$ is block-diagonal with sub-block-bandwidths $(1,1)$;
- $\mathcal{L}^{(a) \rightarrow(a+\tilde{a})}$ is block-diagonal with sub-block-bandwidths $(0,4)$;
- $\mathcal{L}_{W}^{(a) \rightarrow(a-\tilde{a})}$ is block-diagonal with sub-block-bandwidths $(4,0)$;
- $\Delta_{W}^{(1)}$ is block-diagonal with sub-block-bandwidths $(2,2)$.

In order to show the last part of Theorem 3.1, we require the following short lemma.
Lemma 3 For any general parameter $a$ and any $n=0,1, \ldots, k=0, \ldots, n$, we have that

$$
\begin{aligned}
& \frac{\mathrm{d}}{\mathrm{~d} z}\left[w_{R}^{(a+1,2(k+1))} R_{n-k}^{(a, 2 k)}\right] \\
&=w_{R}^{(a+1,2(k+1))} R_{n-k}^{(a, 2 k) \prime \prime}-2(k+1) z w_{R}^{(a+1,2 k)} R_{n-k}^{(a, 2 k) \prime}+(a+1) w_{R}^{(a, 2(k+1))} R_{n-k}^{(a, 2 k) \prime} \\
&=\sum_{m=n-1}^{n+1} c_{m, k} w_{R}^{(a, 2 k)} R_{m-k}^{(a, 2 k)},
\end{aligned}
$$

where

$$
c_{m, k}=-\frac{1}{\omega_{R}^{(a, 2 k)}} \int_{\alpha}^{1} R_{n-k}^{(a, 2 k) \prime} R_{m-k}^{(a, 2 k) \prime \prime} w_{R}^{(a+1,2(k+1))} \mathrm{d} z
$$

Proof of Lemma 3. Since $\frac{\mathrm{d}}{\mathrm{d} z}\left[w_{R}^{(a+1,2(k+1))} R_{n-k}^{(a, 2 k) \prime}\right]=w_{R}^{(a, 2 k)} r_{n-k+1}$ where $r_{n-k+1}$ is a degree $n-k+1$ polynomial, we have that

$$
\frac{\mathrm{d}}{\mathrm{~d} z}\left[w_{R}^{(a+1,2(k+1))} R_{n-k}^{(a, 2 k)}\right]=\sum_{m=0}^{n-k+1} \tilde{c}_{\{n, k\}, m} w_{R}^{(a, 2 k)} R_{m}^{(a, 2 k)}
$$

for some coefficients $\tilde{c}_{\{n, k\}, m}$. These coefficients are given by

$$
\begin{aligned}
\tilde{c}_{\{n, k\}, m} & =\frac{1}{\omega_{R}^{(a, 2 k)}}\left\langle\frac{\mathrm{d}}{\mathrm{~d} z}\left[w_{R}^{(a+1,2(k+1))} R_{n-k}^{(a, 2 k) \prime}\right], R_{m}^{(a, 2 k)}\right\rangle_{w_{R}^{(0,0)}} \\
& =-\frac{1}{\omega_{R}^{(a, 2 k)}} \int_{\alpha}^{1} R_{n-k}^{(a, 2 k)^{\prime}} R_{m}^{(a, 2 k)^{\prime}} w_{R}^{(a+1,2(k+1))} \mathrm{d} z
\end{aligned}
$$

We show that these are zero for $m<n-k-1$ by integrating twice by parts:

$$
\begin{aligned}
& \left\langle\frac{\mathrm{d}}{\mathrm{~d} z}\left[w_{R}^{(a+1,2(k+1))} R_{n-k}^{(a, 2 k) \prime}\right], R_{m}^{(a, 2 k)}\right\rangle_{w_{R}^{(0,0)}} \\
& \quad=-\int_{\alpha}^{1} R_{n-k}^{(a, 2 k) \prime} R_{m-k}^{(a, 2 k) \prime} w_{R}^{(a+1,2(k+1))} \mathrm{d} z \\
& =\int_{\alpha}^{1} R_{n-k}^{(a, 2 k) \prime}\left[(a+1) R_{m}^{(a, 2 k) \prime} w_{R}^{(0,2)}\right. \\
& \left.\quad-2(k+1) z R_{m}^{(a, 2 k) \prime} w_{R}^{(1,0)}+R_{m}^{(a, 2 k) \prime \prime} w_{R}^{(1,2)}\right] w_{R}^{(a, 2 k)} \mathrm{d} z
\end{aligned}
$$

which is indeed zero for $m<n-k-1$ by orthogonality.
Proof of Theorem 3.1. For the operator $D_{\theta}$ for partial differentiation by $\theta$, we simply have that

$$
\begin{aligned}
\frac{\partial}{\partial \theta} Q_{n, k, i}^{(a)}(x, y, z) & =R_{n-k}^{(a, 2 k)}(z) \rho(z)^{k} \frac{\mathrm{~d}}{\mathrm{~d} \theta} Y_{k, i}(\theta) \\
& = \begin{cases}(-1)^{i+1} k Q_{n, k,|i-1|}^{(a)}(x, y, z) & k>0 \\
0 & k=0\end{cases}
\end{aligned}
$$

We now proceed with the case for the operator $D_{\varphi}^{(a)}$ for partial differentiation by $\varphi$. The entries of the operator are given by the coefficients in the expansion

$$
\rho \frac{\partial}{\partial \varphi} Q_{n, k, i}^{(a)}=\sum_{m=0}^{n+1} \sum_{j=0}^{m} \sum_{h=0}^{1} c_{m, j, h} Q_{m, j, h}^{(a+1)}
$$

where the coefficients are

$$
c_{m, j, h}=\left\|Q_{m, j, h}^{(a+1)}\right\|_{Q^{(a+1)}}^{-2}\left\langle\rho \frac{\partial}{\partial \varphi} Q_{n, k, i}^{(a)}, Q_{m, j, h}^{(a+1)}\right\rangle_{Q^{(a+1)}}
$$

Now, note that

$$
\begin{aligned}
w_{R}^{(a, b) \prime}(z) & =a w_{R}^{(a-1, b)}(z)+c \rho(z) \rho^{\prime}(z) w_{R}^{(a, b-2)}(z), \\
\rho(z) \rho^{\prime}(z) & =-z, \\
\frac{\partial}{\partial \varphi} Q_{n, k, i}^{(a)}(x, y, z) & =-\rho(z) \frac{\mathrm{d}}{\mathrm{~d} z}\left[\rho(z)^{k} R_{n-k}^{(a, 2 k)}(z)\right] Y_{k, i}(\theta), \\
\frac{\partial}{\partial \varphi}\left[w_{R}^{(a, 0)}(z) Q_{n, k, i}^{(a)}(x, y, z)\right] & =-\rho(z) \frac{\mathrm{d}}{\mathrm{~d} z}\left[w_{R}^{(a, k)}(z) R_{n-k}^{(a, 2 k)}(z)\right] Y_{k, i}(\theta) .
\end{aligned}
$$

Then,

$$
\begin{aligned}
& \left\langle\rho \frac{\partial}{\partial \varphi} Q_{n, k, i}^{(a)}, Q_{m, j, h}^{(a+1)}\right\rangle_{Q^{(a+1)}} \\
& =-\int_{\alpha}^{1}\left(\int_{0}^{2 \pi} \rho(z)^{2} \frac{\mathrm{~d}}{\mathrm{~d} z}\left[R_{n-k}^{(a, 2 k)}(z) \rho(z)^{k}\right] R_{m-j}^{(a+1,2 j)}(z) \rho(z)^{j} Y_{k, i}(\theta) Y_{j, h}(\theta) \mathrm{d} \theta\right) w_{R}^{(a+1,0)} \mathrm{d} z \\
& =\left(\int_{0}^{2 \pi} Y_{k, i}(\theta) Y_{j, h}(\theta) \mathrm{d} \theta\right)\left(\int_{\alpha}^{1} R_{m-j}^{(a+1,2 j)}\left[k z R_{n-k}^{(a, 2 k)}-\rho^{2} R_{n-k}^{(a, 2 k) \prime}\right] w_{R}^{(a+1, k+j)} \mathrm{d} z\right) \\
& =\pi \delta_{k, j} \delta_{i, h} \int_{\alpha}^{1} R_{m-k}^{(a+1,2 k)}\left[k z R_{n-k}^{(a, 2 k)}-\rho^{2} R_{n-k}^{(a, 2 k) \prime}\right] w_{R}^{(a+1,2 k)} \mathrm{d} z
\end{aligned} \quad \begin{aligned}
& =\pi \delta_{k, j} \delta_{i, h} \int_{\alpha}^{1} R_{n-k}^{(a, 2 k)}\left\{k z R_{m-k}^{(a+1,2 k)} w_{R}^{(1,0)}+R_{m-k}^{(a+1,2 k) \prime} w_{R}^{(1,2)}\right. \\
& \left.\quad+a \rho^{2} R_{m-k}^{(a+1,2 k)}-(2 k+2) z R_{m-k}^{(a+1,2 k)} w_{R}^{(1,0)}\right\} w_{R}^{(a, 2 k)} \mathrm{d} z
\end{aligned}
$$

which is zero for $j \neq k, h \neq i$, and $m<n-2$ by orthogonality.
Similarly, for the operator $W_{\varphi}^{(a)}$ for partial differentiation by $\varphi$ on the weighted space, the entries of the operator are given by the coefficients in the expansion $\rho \frac{\partial}{\partial \varphi}\left(w_{R}^{(a, 0)} Q_{n, k, i}^{(a)}=\right.$ $\sum_{m=0}^{n+2} \sum_{j=0}^{m} \sum_{h=0}^{1} c_{m, j, h} w_{R}^{(a-1,0)} Q_{m, j, h}^{(a-1)}$, where the coefficients are

$$
c_{m, j, h}=\left\|Q_{m, j, h}^{(a-1)}\right\|_{Q^{(a-1)}}^{-2}\left\langle\rho \frac{\partial}{\partial \varphi}\left(w_{R}^{(a, 0)} Q_{n, k, i}^{(a)}\right), Q_{m, j, h}^{(a-1)}\right\rangle_{Q^{(0)}} .
$$

Now,

$$
\begin{aligned}
& \left\langle\rho \frac{\partial}{\partial \varphi}\left(w_{R}^{(a, 0)} Q_{n, k, i}^{(a)}\right), Q_{m, j, h}^{(a-1)}\right\rangle_{Q^{(0)}} \\
& \quad=-\int_{\alpha}^{1}\left(\int_{0}^{2 \pi} \rho(z)^{2} \frac{\mathrm{~d}}{\mathrm{~d} z}\left[R_{n-k}^{(a, 2 k)}(z) w_{R}^{(a, k)}(z)\right] R_{m-j}^{(a-1,2 j)}(z) \rho(z)^{j} Y_{k, i}(\theta) Y_{j, h}(\theta) \mathrm{d} \theta\right) \mathrm{d} z \\
& \quad=\left(\int_{0}^{2 \pi} Y_{k, i}(\theta) Y_{j, h}(\theta) \mathrm{d} \theta\right) \\
& \quad \cdot\left(\int_{\alpha}^{1} R_{m-j}^{(a-1,2 j)}\left[k z R_{n-k}^{(a, 2 k)} w_{R}^{(1,0)}-R_{n-k}^{(a, 2 k) \prime} w_{R}^{(1,2)}-a R_{n-k}^{(a, 2 k)} \rho^{2}\right] w_{R}^{(a-1, k+j)} \mathrm{d} z\right)
\end{aligned}
$$

$$
\begin{aligned}
& =\pi \delta_{k, j} \delta_{i, h} \int_{\alpha}^{1} R_{m-k}^{(a-1,2 k)}\left[k z R_{n-k}^{(a, 2 k)} w_{R}^{(1,0)}-R_{n-k}^{(a, 2 k) \prime} w_{R}^{(1,2)}-a R_{n-k}^{(a, 2 k)} \rho^{2}\right] w_{R}^{(a-1,2 k)} \mathrm{d} z \\
& =\pi \delta_{k, j} \delta_{i, h} \int_{\alpha}^{1} R_{n-k}^{(a, 2 k)}\left\{k z R_{m-k}^{(a-1,2 k)} w_{R}^{(1,0)}-a \rho^{2} R_{m-k}^{(a-1,2 k)}+R_{m-k}^{(a-1,2 k) \prime} w_{R}^{(1,2)}\right. \\
& \left.\quad+a \rho^{2} R_{m-k}^{(a-1,2 k)}-(2 k+2) z R_{m-k}^{(a-1,2 k)} w_{R}^{(1,0)}\right\} w_{R}^{(a-1,2 k)} \mathrm{d} z \\
& =\pi \delta_{k, j} \delta_{i, h} \int_{\alpha}^{1} R_{n-k}^{(a, 2 k)}\left\{k z R_{m-k}^{(a-1,2 k)}+R_{m-k}^{(a-1,2 k) \prime} \rho^{2}-(2 k+2) z R_{m-k}^{(a-1,2 k)}\right\} w_{R}^{(a, 2 k)} \mathrm{d} z,
\end{aligned}
$$

which is zero for $j \neq k, h \neq i$, and $m<n-1$ by orthogonality.
We move on to the spherical Laplacian operators. Note that the Laplacian acting on the weighted and non-weighted spherical cap OP $Q_{n, k, i}^{(a)}$ yield

$$
\begin{align*}
&\left.\begin{array}{rl}
\Delta_{\mathrm{S}} Q_{n, k, i}^{(a)}= & \frac{1}{\rho} \frac{\partial}{\partial \varphi}(\rho
\end{array} \frac{\partial}{\partial \varphi}\left[R_{n-k}^{(a, 2 k)}(\cos \varphi) \sin ^{k} \varphi\right]\right) Y_{k, i}(\theta) \\
&+R_{n-k}^{(a, 2 k)}(\cos \varphi) \sin ^{k-2} \varphi \frac{\partial^{2}}{\partial \theta^{2}} Y_{k, i}(\theta) \\
&= Y_{k, i}(\theta) \rho(z)^{k}\left\{-k(k+1) R_{n-k}^{(a, 2 k)}(z)-2(k+1) z R_{n-k}^{(a, 2 k) \prime}(z)\right. \\
&\left.+\rho(z)^{2} R_{n-k}^{(a, 2 k) \prime \prime}(z)\right\},  \tag{14}\\
& \Delta_{\mathrm{S}}\left(w_{R}^{(a, 0)} Q_{n, k, i}^{(a)}\right)=\frac{1}{\rho} \frac{\partial}{\partial \varphi}\left(\rho \frac{\partial}{\partial \varphi}\left[w_{R}^{(a, 0)}(\cos \varphi) R_{n-k}^{(a, 2 k)}(\cos \varphi) \sin ^{k} \varphi\right]\right) Y_{k, i}(\theta) \\
&+w_{R}^{(a, 0)}(\cos \varphi) R_{n-k}^{(a, 2 k)}(\cos \varphi) \sin ^{k-2} \varphi \frac{\partial^{2}}{\partial \theta^{2}} Y_{k, i}(\theta) \\
&= Y_{k, i}(\theta)\left\{R_{n-k}^{(a, 2 k)}(z)\left[-k(k+1) w_{R}^{(a, k)}(z)-2 a(k+1) z w_{R}^{(a-1, k)}(z)\right]\right. \\
&+a(a-1) R_{n-k}^{(a, 2 k)}(z) w_{R}^{(a-2, k+1)}(z) \\
&+R_{n-k}^{(a, 2 k) \prime}(z)\left[-2(k+1) z w_{R}^{(a, k)}(z)+2 a w_{R}^{(a-1, k+2)}(z)\right] \\
&\left.+R_{n-k}^{(a, 2 k) \prime \prime}(z) w_{R}^{(a, k+2)}(z)\right\} . \tag{15}
\end{align*}
$$

For the operator $\mathcal{L}^{(a) \rightarrow(a+\tilde{a})}$ for the surface Laplacian on a non-weighted space, the entries of the operator are given by the coefficients in the expansion

$$
\Delta_{\mathrm{S}} Q_{n, k, i}^{(a)}=\sum_{m=0}^{n} \sum_{j=0}^{m} \sum_{h=0}^{1} c_{m, j, h} Q_{m, j, h}^{(a+\tilde{a})}
$$

where the coefficients are

$$
c_{m, j, h}=\left\|Q_{m, j, h}^{(a+\tilde{a})}\right\|_{Q^{(a+\tilde{a})}}^{-2}\left\langle\Delta_{\mathrm{S}} Q_{n, k, i}^{(a)}, Q_{m, j, h}^{(a+\tilde{a})}\right\rangle_{Q^{(a+\tilde{a})}}
$$

Using equation (14), and integrating by parts twice, we then have that

$$
\begin{aligned}
& \left\langle\Delta_{\mathrm{S}} Q_{n, k, i}^{(a)}, Q_{m, j}^{(a+\tilde{a})}\right\rangle_{Q^{(a+\tilde{a})}} \\
& =\left(\int_{0}^{2 \pi} Y_{k, i}(\theta) Y_{j, h}(\theta) \mathrm{d} \theta\right) \\
& \quad \cdot\left(\int _ { \alpha } ^ { 1 } R _ { m - j } ^ { ( a + \tilde { a } , 2 j ) } w _ { R } ^ { ( a + \tilde { a } ) , k + j ) } \left\{-k(k+1) R_{n-k}^{(a, 2 k)}-2(k+1) z R_{n-k}^{(a, 2 k) \prime}\right.\right. \\
& \\
& \left.\left.\quad+\rho(z)^{2} R_{n-k}^{(a, 2 k) \prime \prime}\right\} \mathrm{d} z\right)
\end{aligned} \quad \begin{aligned}
& =\pi \delta_{k, j} \delta_{i, h} \int_{\alpha}^{1} R_{m-k}^{(a+\tilde{a}, 2 k)} w_{R}^{(a+\tilde{a}, 0)}\left(-k(k+1) R_{n-k}^{(a, 2 k)} \rho^{2 k}+\frac{\mathrm{d}}{\mathrm{~d} z}\left[R_{n-k}^{(a, 2 k) \prime} \rho^{2(k+1)}\right]\right) \mathrm{d} z \\
& =\pi \delta_{k, j} \delta_{i, h} \int_{\alpha}^{1}\left\{-k(k+1) R_{m-k}^{(a+\tilde{a}, 2 k)} R_{n-k}^{(a, 2 k)} w_{R}^{(a+\tilde{a}, 2 k)}\right. \\
& \left.\quad-R_{n-k}^{(a, 2 k) \prime} w_{R}^{(a+\tilde{a}-1,2 k)}\left[R_{m-k}^{(a+\tilde{a}, 2 k) \prime} w_{R}^{(1,0)}+(a+\tilde{a}) R_{m-k}^{(a+\tilde{a}, 2 k)}\right]\right\} \mathrm{d} z \\
& =\pi \delta_{k, j} \delta_{i, h} \int_{\alpha}^{1} R_{n-k}^{(a, 2 k)} w_{R}^{(a, 2 k)} r_{m-k+\tilde{a}} \mathrm{~d} z,
\end{aligned}
$$

where $r_{m-k+\tilde{a}}$ is a degree $m-k+\tilde{a}$ polynomial in $z$, and so the above is zero for $n-k>m-k+\tilde{a} \Longleftrightarrow$ $m<n-\tilde{a}$.

For the operator $\mathcal{L}_{W}^{(a) \rightarrow(a-\tilde{a})}$ for the surface Laplacian on a weighted space, the entries of the operator are given by the coefficients in the expansion

$$
\Delta_{\mathrm{S}}\left(w_{R}^{(a, 0)} Q_{n, k, i}^{(a)}\right)=\sum_{m=0}^{n} \sum_{j=0}^{m} \sum_{h=0}^{1} c_{m, j, h} w_{R}^{(a-\tilde{a}, 0)} Q_{m, j, h}^{(a-\tilde{a})}
$$

where the coefficients are

$$
c_{m, j, h}=\left\|Q_{m, j, h}^{(a-\tilde{a})}\right\|_{Q^{(a-\tilde{a})}}^{-2}\left\langle\Delta_{\mathrm{S}}\left(w_{R}^{(a, 0)} Q_{n, k, i}^{(a)}\right), Q_{m, j, h}^{(a-\tilde{a})}\right\rangle_{Q^{(0)}}
$$

Using equation (15), and integrating by parts thrice, we then have that

$$
\begin{aligned}
& \left\langle\Delta_{\mathrm{S}}\left(w_{R}^{(a, 0)} Q_{n, k, i}^{(a)}\right), Q_{m, j, h}^{(a-\tilde{a})}\right\rangle_{Q^{(0)}} \\
& =\left(\int_{0}^{2 \pi} Y_{k, i}(\theta) Y_{j, h}(\theta) \mathrm{d} \theta\right) \\
& \cdot\left(\int _ { \alpha } ^ { 1 } R _ { m - j } ^ { ( a - \tilde { a } , 2 j ) } w _ { R } ^ { ( a - 2 , k + j ) } \left\{R_{n-k}^{(a, 2 k)}\left[-k(k+1) w_{R}^{(2,0)}-2 a(k+1) z w_{R}^{(1,0)}+a(a-1) \rho^{2}\right]\right.\right. \\
& +R_{n-k}^{(a, 2 k) \prime}\left[-2(k+1) z w_{R}^{(2,0)}+2 a w_{R}^{(1,2)}\right] \\
& \left.\left.+R_{n-k}^{(a, 2 k) \prime \prime} w_{R}^{(2,2)}\right\} \mathrm{d} z\right) \\
& =\pi \delta_{k, j} \delta_{i, h} \int_{\alpha}^{1}\left\{R_{m-k}^{(a-\tilde{a} 2 k)} R_{n-k}^{(a, 2 k)} w_{R}^{(a-2,2 k)}\left[-k(k+1) w_{R}^{(2,0)}-2 a(k+1) z w_{R}^{(1,0)}+a(a-1) \rho^{2}\right]\right. \\
& +a R_{n-k}^{(a, 2 k) \prime} R_{m-k}^{(a-\tilde{a}, 2 k)} w_{R}^{(a-1,2 k+2)} \\
& \left.+R_{m-k}^{(a-\tilde{a}, 2 k)} \frac{\mathrm{d}}{\mathrm{~d} z}\left[R_{n-k}^{(a, 2 k) \prime} w_{R}^{(a, 2 k+2)}\right]\right\} \mathrm{d} z \\
& =\pi \delta_{k, j} \delta_{i, h} \int_{\alpha}^{1}\left\{R_{m-k}^{(a-\tilde{a} 2 k)} R_{n-k}^{(a, 2 k)} w_{R}^{(a-2,2 k)}\left[-k(k+1) w_{R}^{(2,0)}-2 a(k+1) z w_{R}^{(1,0)}+a(a-1) \rho^{2}\right]\right. \\
& +a R_{n-k}^{(a, 2 k) \prime} R_{m-k}^{(a-\tilde{a}, 2 k)} w_{R}^{(a-1,2 k+2)} \\
& \left.+R_{n-k}^{(a, 2 k)} w_{R}^{(a-1,2 k)}\left[R_{m-k}^{(a-\tilde{a}, 2 k) \prime \prime} w_{R}^{(1,2)}+a R_{m-k}^{(a-\tilde{a}, 2 k) \prime} \rho^{2}-2(k+1) z R_{m-k}^{(a-\tilde{a}, 2 k)} w_{R}^{(1,0)}\right]\right\} \mathrm{d} z \\
& =\pi \delta_{k, j} \delta_{i, h} \int_{\alpha}^{1}\left\{R_{m-k}^{(a-\tilde{a}, 2 k)} R_{n-k}^{(a, 2 k)} w_{R}^{(a-2,2 k)}\left[-k(k+1) w_{R}^{(2,0)}-2 a(k+1) z w_{R}^{(1,0)}+a(a-1) \rho^{2}\right]\right. \\
& +R_{n-k}^{(a, 2 k)} w_{R}^{(a-1,2 k+2)}\left[R_{m-k}^{(a-\tilde{a}, 2 k) \prime \prime} \rho^{2}-2(k+1) z R_{m-k}^{(a-\tilde{a}, 2 k) \prime}\right] \\
& \left.+a\left[R_{n-k}^{(a, 2 k)} R_{m-k}^{(a-\tilde{a}, 2 k) \prime}+R_{m-k}^{(a-\tilde{a}, 2 k)} R_{n-k}^{(a, 2 k) \prime}\right] w_{R}^{(a-1,2 k+2)}\right\} \mathrm{d} z \\
& =\pi \delta_{k, j} \delta_{i, h} \int_{\alpha}^{1}\left\{R_{m-k}^{(a-\tilde{a}, 2 k)} R_{n-k}^{(a, 2 k)} w_{R}^{(a-2,2 k)}\left[-k(k+1) w_{R}^{(2,0)}-2 a(k+1) z w_{R}^{(1,0)}+a(a-1) \rho^{2}\right]\right. \\
& +R_{n-k}^{(a, 2 k)} w_{R}^{(a-1,2 k+2)}\left[R_{m-k}^{(a-\tilde{a}, 2 k) \prime \prime} \rho^{2}-2(k+1) z R_{m-k}^{(a-\tilde{a}, 2 k) \prime}\right] \\
& \left.-a R_{n-k}^{(a, 2 k)} R_{m-k}^{(a-\tilde{a}, 2 k)} w_{R}^{(a-2,2 k)}\left[(a-1) \rho^{2}-2(k+1) z w_{R}^{(1,0)}\right]\right\} \mathrm{d} z \\
& =\pi \delta_{k, j} \delta_{i, h} \int_{\alpha}^{1} R_{n-k}^{(a, 2 k)} w_{R}^{(a, 2 k)} r_{m-k} \mathrm{~d} z,
\end{aligned}
$$

where $r_{m-k}$ is a degree $m-k$ polynomial in $z$, and so the above is zero for $n-k>m-k \Longleftrightarrow m<n$.
Finally, fix $a=1$. For the operator $\Delta_{W}^{(1)}$ for the Laplacian on the weighted space, the entries of the operator are given by the coefficients in the expansion $\Delta_{\mathrm{S}}\left(w_{R}^{(1,0)} Q_{n, k, i}^{(1)}\right)=$
$\sum_{m=0}^{n+2} \sum_{j=0}^{m} \sum_{h=0}^{1} c_{m, j, h} Q_{m, j, h}^{(1)}$, where the coefficients are given by

$$
c_{m, j, h}=\left\|Q_{m, j, h}^{(1)}\right\|_{Q^{(1)}}^{-2}\left\langle\Delta_{\mathrm{S}}\left(w_{R}^{(1,0)} Q_{n, k, i}^{(1)}\right), Q_{m, j, h}^{(1)}\right\rangle Q^{(1)}
$$

Using equation (15) with $a=1$, and Lemma 3, we then have that

$$
\begin{aligned}
& \left\langle\Delta_{\mathrm{S}}\left(w_{R}^{(1,0)} Q_{n, k, i}^{(1)}\right), Q_{m, j, h}^{(1)}\right\rangle_{Q^{(1)}} \\
& =\left(\int_{0}^{2 \pi} Y_{k, i}(\theta) Y_{j, h}(\theta) \mathrm{d} \theta\right) \\
& \cdot\left(\int _ { \alpha } ^ { 1 } R _ { m - j } ^ { ( 1 , 2 j ) } \left\{R_{n-k}^{(1,2 k)}\left[-k^{2} w_{R}^{(1, k)}-w_{R}^{(1, k)}-2(k+1) z w_{R}^{(0, k)}\right]\right.\right. \\
& +R_{n-k}^{(1,2 k) \prime}\left[2 w_{R}^{(0, k+2)}-2(k+1) z w_{R}^{(1, k)}\right] \\
& \left.\left.+R_{n-k}^{(1,2 k) \prime \prime} w_{R}^{(1, k+2)}\right\} w_{R}^{(1, j)} \mathrm{d} z\right) \\
& =\pi \delta_{k, j} \delta_{i, h} \int_{\alpha}^{1} R_{m-k}^{(1,2 k)}\left\{R_{n-k}^{(1,2 k)}\left[-k(k+1) w_{R}^{(1,0)}-2(k+1) z+c_{n, k}\right]\right. \\
& \left.+c_{n-1, k} R_{n-k-1}^{(1,2 k)}+c_{n+1, k} R_{n-k+1}^{(1,2 k)}\right\} w_{R}^{(1,2 k)} \mathrm{d} z \\
& =-\pi \delta_{k, j} \delta_{i, h}\left(\delta_{m, n-1}+\delta_{m, n}+\delta_{m, n+1}\right) \int_{\alpha}^{1}\left\{R_{n-k}^{(1,2 k)} R_{m-k}^{(1,2 k)}\left(k(k+1) w_{R}^{(1,0)}+2(k+1) z\right)\right. \\
& \left.+R_{n-k}^{(1,2 k) \prime} R_{m-k}^{(1,2 k) \prime} w_{R}^{(2,2(k+1))}\right\} \mathrm{d} z,
\end{aligned}
$$

where the $c_{n-1, k}, c_{n, k}, c_{n+1, k}$ are those derived in Lemma 3.
By applying these differential operators, we are (in some cases) incrementing or decrementing the parameter value $a$. It is therefore necessary to also be able to raise or lower the parameter by way of an independent operator. There exist conversion operators that do exactly this, transforming the OPs from one (weighted or non-weighted) parameter space to another.
Definition 7 Define the operator matrices $T^{(a) \rightarrow(a+\tilde{a})}, \quad T_{W}^{(a) \rightarrow(a-\tilde{a})}$ for conversion between nonweighted spaces and weighted spaces, respectively, according to

$$
\begin{aligned}
\tilde{\mathbb{Q}}_{N}^{(a)}(x, y, z) & =\left(T^{(a) \rightarrow(a+\tilde{a})}\right)^{\top} \tilde{\mathbb{Q}}_{N}^{(a+\tilde{a})}(x, y, z) \\
\mathbb{W}_{N}^{(a)}(x, y, z) & =\left(T_{W}^{(a) \rightarrow(a-\tilde{a})}\right)^{\top} \mathbb{W}_{N}^{(a-\tilde{a})}(x, y, z) .
\end{aligned}
$$

Lemma 4 The operator matrices in Definition 7 are sparse, with banded-block-banded structure. More specifically,

- $T^{(a) \rightarrow(a+\tilde{a})}$ is block-diagonal with sub-block bandwidths $(0,2 \tilde{a})$;
- $T_{W}^{(a) \rightarrow(a-\tilde{a})}$ is block-diagonal with sub-block bandwidths (2 $2 \tilde{a}, 0$ ).

Proof. We proceed with the case for the non-weighted operators $T^{(a) \rightarrow(a+\tilde{a})}$. Since $\left\{Q_{m, j, h}^{(a+\tilde{a})}\right\}$ for $m=0, \ldots, n, j=0, \ldots, m, h=0,1$ is an orthogonal basis for any degree $n$ polynomial, we can expand $Q_{n, k, i}^{(a)}=\sum_{m=0}^{n} \sum_{j=0}^{m} t_{m, j} Q_{m, j, h}^{(a+\tilde{a})}$. The coefficients of the expansion are then the entries of the operator matrix. We will show that the only non-zero coefficients are for $k=j, i=h$ and $m \geq n-\tilde{a}$. Note that

$$
t_{m, j}=\left\|Q_{m, j, h}^{(a+\tilde{a})}\right\|_{Q^{(a+\tilde{a})}}^{-2}\left\langle Q_{n, k, i}^{(a)}, Q_{m, j, h}^{(a+\tilde{a})}\right\rangle_{Q^{(a+\tilde{a})}},
$$

where

$$
\begin{aligned}
\left\langle Q_{n, k, i}^{(a)}, Q_{m, j, h}^{(a+\tilde{a})}\right\rangle_{Q^{(a+\tilde{a})}} & =\left(\int_{0}^{2 \pi} Y_{k, i}(\theta) Y_{j, h}(\theta) \mathrm{d} \theta\right) \cdot\left(\int_{\alpha}^{1} R_{n-k}^{(a, 2 k)} R_{m-j}^{(a+\tilde{a}, 2 j)} \rho^{k+j} w_{R}^{(a+\tilde{a}, 0)} \mathrm{d} z\right) \\
& =\pi \delta_{k, j} \delta_{i, h} \int_{\alpha}^{1} R_{n-k}^{(a, 2 k)} R_{m-k}^{(a+\tilde{a}, 2 k)} w_{R}^{(a+\tilde{a}, 2 k)} \mathrm{d} z
\end{aligned}
$$

which is zero for $n>m+\tilde{a} \Longleftrightarrow m<n-\tilde{a}$. The sparsity argument for the weighted parameter transformation operator follows similarly.

### 3.1 Further partial differential operators

General linear partial differential operators with polynomial variable coefficients can be constructed by composing the sparse representations for partial derivatives, conversion between bases and Jacobi operators. As a canonical example, we can obtain the matrix operator for the $\rho^{2}$-factored spherical Laplacian $\rho(z)^{2} \Delta_{\mathrm{S}}$ that will take us from coefficients for expansion in the weighted space $\mathbb{W}_{N}^{(1)}(x, y, z)=w_{R}^{(1,0)}(z) \widetilde{\mathbb{Q}}_{N}^{(1)}(x, y, z)$ to coefficients in the non-weighted space $\widetilde{\mathbb{Q}}_{N}^{(1)}(x, y, z)$. Note that this construction will ensure the imposition of the Dirichlet zero boundary conditions on $\Omega$, similar to how the Dirichlet zero boundary conditions would be imposed for the operator $\Delta_{W}^{(1)}$ in Definition 6. The matrix operator for this $\rho^{2}$-factored spherical Laplacian acting on the coefficients vector is then given by

$$
D_{\varphi}^{(0)} W_{\varphi}^{(1)}+T^{(0) \rightarrow(1)} T_{W}^{(1) \rightarrow(0)}\left(D_{\theta}\right)^{2} .
$$

Importantly, this operator will have banded-block-banded structure, and hence will be sparse, as seen in Fig. 1.

Another desirable operator is the Biharmonic operator $\Delta_{\mathrm{S}}^{2}$, for which we assume zero Dirichlet and Neumann conditions. That is,

$$
u(x, y, z)=0, \quad \frac{\partial u}{\partial n}(x, y, z)=\nabla_{S} u(x, y, z) \cdot \hat{\boldsymbol{n}}(x, y, z)=0 \quad \text { for }(x, y, z) \in \partial \Omega
$$

where $\partial \Omega$ is the $z=\alpha$ boundary and $\hat{\boldsymbol{n}}(x, y, z)$ is the outward unit normal vector at the point $(x, y, z)$ on the boundary, i.e., $\hat{\boldsymbol{n}}(x, y, z)=\hat{\boldsymbol{n}}(\boldsymbol{x}):=\frac{\boldsymbol{x}}{\|\boldsymbol{x}\|}=\boldsymbol{x}$. The matrix operator for the Biharmonic operator will take us from coefficients in the space $\mathbb{W}^{(2)}(x, y, z)$ to coefficients in the space $\tilde{\mathbb{Q}}_{N}^{(2)}(x, y, z)$. To construct


Fig. 1. 'Spy' plots of (differential) operator matrices, showing their sparsity. For (c), the weighted variable coefficient Helmholtz operator is $\Delta_{W}^{(1)}+k^{2} T^{(0) \rightarrow(1)} V\left(J_{x}^{(0)^{\top}}, J_{y}^{(0)^{\top}}, J_{z}^{(0)^{\top}}\right) T_{W}^{(1) \rightarrow(0)}$ for $v(x, y, z)=1-\left(3\left(x-x_{0}\right)^{2}+5\left(y-y_{0}\right)^{2}+2\left(z-z_{0}\right)^{2}\right)$ where $\left(x_{0}, z_{0}\right):=(0.7,0.2), y_{0}:=\sqrt{1-x_{0}^{2}-z_{0}^{2}}$ and $k=200$.
this, we can simply multiply together two of the spherical Laplacian operators defined in Definition 6, namely $\mathcal{L}^{(0) \rightarrow(2)}$ and $\mathcal{L}_{W}^{(2) \rightarrow(0)}$ :

$$
\mathcal{B}_{W}^{(2)}:=\mathcal{L}^{(0) \rightarrow(2)} \mathcal{L}_{W}^{(2) \rightarrow(0)}
$$

Since the operator $\mathcal{L}_{W}^{(2) \rightarrow(0)}$ acts on coefficients in the $\mathbb{W}^{(2)}(x, y, z)$ space, we ensure that we satisfy the zero Dirichlet and Neumann boundary conditions-such a function could be written $u(x, y, z)=$ $w_{R}^{(2,0)}(z) \tilde{u}(x, y, z)$ and thus its spherical gradient would be zero on the boundary $z=\alpha$. This allows us to apply the $\mathcal{L}^{(0) \rightarrow(2)}$ operator after, safe in the knowledge that boundary conditions have been accounted for. The sparsity and structure of this biharmonic operator are seen in Fig. 1.

### 3.2 Stability of the Laplacian operator

Denote the condition number of the matrix $A$ :

$$
\kappa(A):=\|A\|_{2}\left\|A^{-1}\right\|_{2},
$$

which encodes how accurate the solution to $A x=b$ is under perturbation. Since the Laplacian operator is block-diagonal one can determine the condition number from the condition numbers of each block on the diagonal, and hence deduce the accuracy in solving the Poisson equation, where graphical representations are seen in Fig. 2 for the $\Delta_{W}^{(1)}$ operator matrix. As the degree $N$ increases, the condition numbers of these blocks becomes grows algebraically fast (at least, the condition number for the first block where $m=0$ does, which will be the largest for each $N$ ). Fortunately, however, by applying a trivial diagonal preconditioning matrix similar to that seen for the ultraspherical method (Olver \& Townsend, 2013), we can significantly bound this growth of the condition numbers-this is also seen in Fig. 2. The preconditioning matrix chosen is $P^{-1}$, where $P$ is the matrix of the diagonal of $\Delta_{W}^{(1)}$.

## 4. Computational aspects

In this section, we discuss how to expand and evaluate functions in our proposed basis and take advantage of the sparsity structure in partial differential operators in practical computational applications.

### 4.1 Constructing $R_{n}^{(a, b)}(x)$

It is possible to recursively obtain the recurrence coefficients for the $\left\{R_{n}^{(a, b)}\right\}$ OPs in (6), see Snowball \& Olver (2020), by careful application of the Christoffel-Darboux formula (Olver et al., 2010, 18.2.12).

### 4.2 Quadrature rule on the spherical cap

In this section, we construct a quadrature rule that is exact for polynomials on the spherical cap $\Omega$ that can be used to expand functions in the OPs $Q_{n, k, i}^{(a)}(x, y, z)$ for a given parameter $a$.
Theorem 4.1 Let $M_{1}, M_{2} \in \mathbb{N}$, and denote the $M_{1}$ Gauss quadrature nodes and weights on $[\alpha, 1]$ with weight $(t-\alpha)^{a}$ as $\left(t_{j}, w_{j}^{(t)}\right)$. Further, denote the $M_{2}$ Gauss quadrature nodes and weights $[-1,1]$ with weight $\left(1-x^{2}\right)^{-\frac{1}{2}}$ as $\left(s_{j}, w_{j}^{(s)}\right)$. Define for $j=1, \ldots, M_{1}, l=1, \ldots, M_{2}$,

$$
\begin{aligned}
x_{l+(j-1) M_{2}} & :=\rho\left(t_{j}\right) s_{l}, \\
y_{l+(j-1) M_{2}} & :=\rho\left(t_{j}\right) \sqrt{1-s_{l}^{2}}, \\
z_{l+(j-1) M_{2}} & :=t_{j}, \\
w_{l+(j-1) M_{2}} & :=w_{j}^{(t)} w_{l}^{(s)} .
\end{aligned}
$$



Fig. 2. Plots of the condition number for the Laplacian operator matrix $\Delta_{W}^{(1)}$ and the preconditioned matrix $P^{-1} \Delta_{W}^{(1)}$ where $P$ is the matrix of the diagonal of $\Delta_{W}^{(1)}$. Top: the condition numbers of each diagonal block of $\Delta_{W}^{(1)}$ for $N=200$, as well as those for $P^{-1} \Delta_{W}^{(1)}$. Bottom: the maximum condition number of all diagonal blocks of the Laplacian and the preconditioned Laplacian operators, as the degree $N$ increases.

Let $f(x, y, z)$ be a function on $\Omega$, and $N \in \mathbb{N}$. The quadrature rule is then

$$
\int_{\Omega} f(x, y, z) w_{R}^{(a, 0)}(z) \mathrm{d} A \approx \sum_{j=1}^{M} w_{j}\left[f\left(x_{j}, y_{j}, z_{j}\right)+f\left(-x_{j},-y_{j}, z_{j}\right)\right],
$$

where $M=M_{1} M_{2}$ and the quadrature rule is exact if $f(x, y, z)$ is a polynomial of degree $\leq N$ with $M_{1} \geq \frac{1}{2}(N+1), M_{2} \geq N+1$.
Remark 3 Note that the Gauss quadrature nodes and weights $\left(t_{j}, w_{j}^{(t)}\right.$ ) will have to be calculated; however, the Gauss quadrature nodes and weights $\left(s_{j}, w_{j}^{(s)}\right)$ are simply the Chebyshev-Gauss quadrature nodes and weights given explicitly (Olver et al., 2010, 3.5.23) as $s_{j}:=\cos \left(\frac{2 j-1}{2 M_{2}} \pi\right), w_{j}^{(s)}:=\frac{\pi}{M_{2}}$.

Proof. Let $f: \Omega \rightarrow \mathbb{R}$. Define the functions $f_{e}, f_{o}: \Omega \rightarrow \mathbb{R}$ by

$$
\begin{array}{ll}
f_{e}(x, y, z):=\frac{1}{2}(f(x, y, z)+f(-x,-y, z)), & \forall(x, y, z) \in \Omega \\
f_{o}(x, y, z):=\frac{1}{2}(f(x, y, z)-f(-x,-y, z)), & \forall(x, y, z) \in \Omega
\end{array}
$$

so that $\mathbf{x} \mapsto f_{e}(\mathbf{x}, z)$ for fixed $z$ is an even function, and $\mathbf{x} \mapsto f_{o}(\mathbf{x}, z)$ for fixed $z$ is an odd function. Note that if $f$ is a polynomial, then $f_{e}(\rho(t) x, \rho(t) y, t)$ is a polynomial in $t \in[\alpha, 1]$ for fixed $(x, y) \in \mathbb{R}^{2}$.

Firstly, we note that

$$
\begin{aligned}
\int_{0}^{2 \pi} g(\cos \theta, \sin \theta) \mathrm{d} \theta & =\int_{-1}^{1}\left[g\left(x, \sqrt{1-x^{2}}\right)+g\left(x,-\sqrt{1-x^{2}}\right)\right] \frac{\mathrm{d} x}{\sqrt{1-x^{2}}} \\
& =\int_{-1}^{1}\left[g\left(x, \sqrt{1-x^{2}}\right)+g\left(-x,-\sqrt{1-x^{2}}\right)\right] \frac{\mathrm{d} x}{\sqrt{1-x^{2}}}
\end{aligned}
$$

for any function $g$, using a change of variables $x \rightarrow-x$ for the second term in the integral. Then, integrating the even function $f_{e}$, we have

$$
\begin{align*}
& \int_{\Omega} f_{e}(x, y, z) w_{R}^{(a, 0)}(z) \mathrm{d} A \\
& =\int_{\alpha}^{1} w_{R}^{(a, 0)}(z)\left(\int_{0}^{2 \pi} f_{e}(\rho(z) \cos \theta, \rho(z) \sin \theta, z) \mathrm{d} \theta\right) \mathrm{d} z \\
& =\int_{\alpha}^{1} w_{R}^{(a, 0)}(z)\left(\int _ { - 1 } ^ { 1 } \left[f_{e}\left(\rho(z) x, \rho(z) \sqrt{1-x^{2}}, z\right)\right.\right. \\
& \left.\left.+f_{e}\left(-\rho(z) x,-\rho(z) \sqrt{1-x^{2}}, z\right)\right] \frac{\mathrm{d} x}{\sqrt{1-x^{2}}}\right) \mathrm{d} z \\
& =2 \int_{\alpha}^{1} w_{R}^{(a, 0)}(z)\left(\int_{-1}^{1} f_{e}\left(\rho(z) x, \rho(z) \sqrt{1-x^{2}}, z\right) \frac{\mathrm{d} x}{\sqrt{1-x^{2}}}\right) \mathrm{d} z \\
& \approx \int_{\alpha}^{1} w_{R}^{(a, 0)}(z)\left(\sum_{l=1}^{M_{2}} w_{l}^{(s)} f_{e}\left(\rho(z) s_{l}, \rho(z) \sqrt{1-s_{l}^{2}}, z\right)\right) \mathrm{d} z \\
& \approx \sum_{j=1}^{M_{1}} w_{j}^{(t)} \sum_{l=1}^{M_{2}} w_{l}^{(s)} f_{e}\left(\rho\left(t_{j}\right) s_{l}, \rho\left(t_{j}\right) \sqrt{1-s_{l}^{2}}, t_{j}\right) \\
& =\sum_{k=1}^{M_{1} M_{2}} w_{j} f_{e}\left(x_{j}, y_{j}, z_{j}\right) .
\end{align*}
$$

Suppose $f$ is a polynomial in $x, y, z$ of degree $N$ and hence that $f_{e}$ is a degree $\leq N$ polynomial. It follows that $s \mapsto f_{e}\left(\rho(z) s, \rho(z) \sqrt{1-s^{2}}, z\right)$ for fixed $z$ is then a polynomial of degree $\leq N$. We therefore achieve equality at $(\star)$ if $2 M_{2}-1 \geq N$ and we achieve equality at ( $\star \star$ ) if also $2 M_{1}-1 \geq N$.

Integrating the odd function $f_{o}$ results in

$$
\begin{aligned}
& \int_{\Omega} f_{o}(x, y, z) w_{R}^{(a, 0)}(z) \mathrm{d} A \\
& \begin{array}{l}
\left.=\int_{\alpha}^{1} w_{R}^{(a, 0)}(z)\left(\int_{0}^{2 \pi} f_{o}(\rho(z) \cos \theta, \rho(z) \sin \theta, z)\right) \mathrm{d} \theta\right) \mathrm{d} z \\
= \\
=\int_{\alpha}^{1} w_{R}^{(a, 0)}(z)\left(\int _ { - 1 } ^ { 1 } \left[f_{o}\left(\rho(z) x, \rho(z) \sqrt{1-x^{2}}, z\right)\right.\right. \\
\left.\left.\quad+f_{o}\left(-\rho(z) x,-\rho(z) \sqrt{1-x^{2}}, z\right)\right] \frac{\mathrm{d} x}{\sqrt{1-x^{2}}}\right) \mathrm{d} z \\
\quad=0
\end{array} \\
& \quad
\end{aligned}
$$

since $f_{o}(x, y, z)=-f_{o}(-x,-y, z)$. Hence, for a polynomial $f$ in $x, y, z$ of degree $N$,

$$
\begin{aligned}
\int_{\Omega} f(x, y, z) w_{R}^{(a, 0)}(z) \mathrm{d} A & =\int_{\Omega}\left(f_{e}(x, y, z)+f_{o}(x, y, z)\right) w_{R}^{(a, 0)}(z) \mathrm{d} A \\
& =\int_{\Omega} f_{e}(x, y, z) w_{R}^{(a, 0)}(z) \mathrm{d} A \\
& =\sum_{j=1}^{M} w_{j} f_{e}\left(x_{j}, y_{j}, z_{j}\right),
\end{aligned}
$$

where $M=M_{1} M_{2}$ and $2 M_{1}-1 \geq N, 2 M_{2}-1 \geq N$.

### 4.3 Obtaining the coefficients for expansion of a function on the spherical cap

Fix $a \in \mathbb{R}$. Then, for any function $f: \Omega \rightarrow \mathbb{R}$, we can express $f$ by

$$
f(x, y, z) \approx \sum_{k=0}^{N} \tilde{\mathbb{Q}}_{N, k}^{(a)}(x, y, z)^{\top} \boldsymbol{f}_{k}=\tilde{\mathbb{Q}}_{N}^{(a)}(x, y, z)^{\top} \boldsymbol{f}
$$

for N sufficiently large, where $\tilde{\mathbb{Q}}_{N, k}^{(a)}, \tilde{\mathbb{Q}}_{N}^{(a)}$ is defined in equations $(11,12,13)$ and where

$$
\begin{aligned}
\boldsymbol{f}_{k} & :=\left(\begin{array}{c}
f_{k, k, 0} \\
f_{k, k, 1} \\
\vdots \\
f_{N, k, 0} \\
f_{N, k, 1}
\end{array}\right) \in \mathbb{R}^{2(N-k+1)} \quad \text { for } n=1,2, \ldots, N, \quad \boldsymbol{f}_{0}:=\left(\begin{array}{c}
f_{0,0,0} \\
\vdots \\
f_{N, 0,0}
\end{array}\right) \in \mathbb{R}^{N+1}, \\
\boldsymbol{f} & :=\left(\begin{array}{c}
f_{0} \\
\vdots \\
f_{N}
\end{array}\right) \in \mathbb{R}^{2(N+1)^{2}}, \quad f_{n, k, i}:=\left\langle f, Q_{n, k, i}^{(a)}\right\rangle_{Q^{(a)}}\left\|Q_{n, k, i}^{(a)}\right\|_{Q^{(a)}}^{-2}
\end{aligned}
$$

Recall from equation (5) that $\left\|Q_{n, k, i}^{(a)}\right\|_{Q^{(a)}}^{2}=\omega_{R}^{(a, 2 k)} \pi$. Using the quadrature rule detailed in Section 4.2 for the inner product, we can calculate the coefficients $f_{n, k, i}$ for each $n=0, \ldots, N, k=0, \ldots, n, i=0,1$ :

$$
\begin{aligned}
f_{n, k, i} & =\frac{1}{2 \omega_{R}^{(a, 2 k)} \pi} \sum_{j=1}^{M} w_{j}\left[f\left(x_{j}, y_{j}, z_{j}\right) Q_{n, k, i}^{(a)}\left(x_{j}, y_{j}, z_{j}\right)+f\left(-x_{j},-y_{j}, z_{j}\right) Q_{n, k, i}^{(a)}\left(-x_{j},-y_{j}, z_{j}\right)\right] \\
& =\frac{1}{M_{2} \omega_{R}^{(a, 2 k)}} \sum_{j=1}^{M}\left[f\left(x_{j}, y_{j}, z_{j}\right) Q_{n, k, i}^{(a)}\left(x_{j}, y_{j}, z_{j}\right)+f\left(-x_{j},-y_{j}, z_{j}\right) Q_{n, k, i}^{(a)}\left(-x_{j},-y_{j}, z_{j}\right)\right]
\end{aligned}
$$

where the quadrature nodes and weights are those from Theorem 4.1 , and $M=M_{1} M_{2}$ with $2 M_{1}-1 \geq$ $N, M_{2}-1 \geq N$ (i.e., we can choose $M_{2}:=N+1$ and $M_{1}:=\left\lceil\frac{N+1}{2}\right\rceil$ ).
Remark 4 While it may be possible to leverage fast transforms, as in Slevinsky (2018, 2019), to speed up the calculation of the coefficients, this remains an open question.

### 4.4 Function evaluation

For a function $f$, with coefficients vector $\boldsymbol{f}$ for expansion in the $\left\{Q_{n, k, i}\right\}$ basis as determined via the method in Section 4.3 up to order $N$, we can use the Clenshaw algorithm to evaluate the function at a point $(x, y, z) \in \Omega$ as follows. Let $A_{n}, B_{n}, D_{n}^{\top}, C_{n}$ be the Clenshaw matrices from Definition 5, and
define the rearranged coefficients vector $\tilde{f}$ via

$$
\begin{aligned}
\boldsymbol{f}_{n} & :\left(\begin{array}{c}
f_{n, 0,0} \\
f_{n, 1,0} \\
f_{n, 1,1} \\
\vdots \\
f_{n, n, 0} \\
f_{n, n, 1}
\end{array}\right) \in \mathbb{R}^{2(N+1)} \quad \text { for } n=1,2, \ldots, N, \quad f_{0}=f_{0,0,0} \in \mathbb{R} \\
\boldsymbol{f} & :=\left(\begin{array}{c}
f_{0} \\
\vdots \\
\boldsymbol{f}_{N}
\end{array}\right) \in \mathbb{R}^{(N+1)^{2}} .
\end{aligned}
$$

The trivariate Clenshaw algorithm works similar to the bivariate Clenshaw algorithm introduced in Olver et al. (2019) for expansions in the triangle:

$$
\begin{aligned}
& \text { (1) set } \boldsymbol{\xi}_{N+2}=\mathbf{0}, \boldsymbol{\xi}_{N+2}=\mathbf{0} \\
& \text { (2) for } n=N:-1: 0 \\
& \quad \operatorname{set} \boldsymbol{\xi}_{n}^{T}=f_{n}^{T}-\boldsymbol{\xi}_{n+1}^{T} D_{n}^{T}\left(B_{n}-G_{n}(x, y, z)\right)-\xi_{n+2}^{T} D_{n+1}^{T} C_{n+1}
\end{aligned}
$$

(3) output: $f(x, y, z) \approx \boldsymbol{\xi}_{0} \mathbb{Q}_{0}^{(a)}=\xi_{0} Q_{0}^{(a)}$.

### 4.5 Calculating non-zero entries of the operator matrices

The proofs of Theorem 3.1 and Lemma 4 provide a way to calculate the non-zero entries of the operator matrices given in Definitions 6 and 7. We can simply use quadrature to calculate the one-dimensional inner products, which has a complexity of $\mathcal{O}\left(N^{2}\right)$. This proves much cheaper computationally than using the three-dimensional quadrature rule to calculate the surface inner products, which has a complexity of $\mathcal{O}\left(N^{3}\right)$.

### 4.6 Obtaining operator matrices for variable coefficients

The Clenshaw algorithm outlined in Section 4.4 can also be used with Jacobi matrices $J_{x}^{(a)}, J_{y}^{(a)}, J_{z}^{(a)}$ replacing the point $(x, y, z)$. Let $v: \Omega \rightarrow \mathbb{R}$ be the function that we wish to obtain an operator matrix $V$ for $v$, so that

$$
v(x, y, z) f(x, y, z)=v(x, y, z) \boldsymbol{f} \mathbb{Q}^{(a)}(x, y, z)=(V \boldsymbol{f})^{\top} \mathbb{Q}^{(a)}(x, y, z),
$$

i.e., $V \boldsymbol{f}$ is the coefficients vector for the function $v(x, y, z) f(x, y, z)$.

To this end, let $\tilde{\boldsymbol{v}}$ be the coefficients for expansion up to order $N$ in the $\left\{Q_{n, k, i}\right\}$ basis of $v$ (rearranged as in Section 4.4 so that $\left.v(x, y, z)=\boldsymbol{v}^{\top} \mathbb{Q}^{(a)}(x, y, z)\right)$. Denote $X:=\left(J_{x}^{(a)}\right)^{\top}, Y:=\left(J_{y}^{(a)}\right)^{\top}, Z:=\left(J_{z}^{(a)}\right)^{\top}$.

The operator $V$ is then the result of the following:

$$
\begin{aligned}
& \text { (1) set } \boldsymbol{\xi}_{N+2}=\mathbf{0}, \boldsymbol{\xi}_{N+2}=\mathbf{0} \\
& \text { (2) for } n=N:-1: 0 \\
& \qquad \text { set } \xi_{n}^{T}=\tilde{\boldsymbol{v}}_{n}^{T}-\boldsymbol{\xi}_{n+1}^{T} D_{n}^{T}\left(B_{n}-G_{n}(X, Y, Z)\right)-\boldsymbol{\xi}_{n+2}^{T} D_{n+1}^{T} C_{n+1} \\
& \text { (3) output: } V=v(X, Y, Z) \approx \xi_{0} \mathbb{Q}_{0}^{(a)}=\xi_{0} Q_{0}^{(a)}
\end{aligned}
$$

where at each iteration, $\boldsymbol{\xi}_{n}$ is a vector of matrices.

## 5. Examples on spherical caps with zero Dirichlet conditions

We now demonstrate how the sparse linear systems constructed as above can be used to efficiently solve PDEs with zero Dirichlet conditions on the spherical cap defined by $\Omega$. We consider Poisson, inhomogeneous variable coefficient Helmholtz equation and the Biharmonic equation, as well as a time dependent heat equation, demonstrating the versatility of the approach.

### 5.1 Poisson

The Poisson equation is the classic problem of finding $u(x, y, z)$ given a function $f(x, y, z)$ such that

$$
\begin{cases}\Delta_{\mathrm{S}} u(x, y, z)=f(x, y, z) & \text { in } \Omega  \tag{16}\\ u(x, y, z)=0 & \text { on } \partial \Omega\end{cases}
$$

noting the imposition of zero Dirichlet boundary conditions on $u$.
We can tackle the problem as follows. Choose an $N \in \mathbb{N}$ large enough for the problem, and denote the coefficient vector for expansion of $u$ in the $\mathbb{W}_{N}^{(1)}$ OP basis up to degree $N$ by $\mathbf{u}$, and the coefficient vector for expansion of $f$ in the $\widetilde{\mathbb{Q}}_{N}^{(1)}$ OP basis up to degree $N$ by $\mathbf{f}$. Since $f$ is known, we can obtain $\mathbf{f}$ using the quadrature rule in Section 4.3. In matrix-vector notation, our system hence becomes

$$
\Delta_{W}^{(1)} \mathbf{u}=\mathbf{f},
$$

which can be solved to find u. In Fig. 3, we see the solution to the Poisson equation with zero boundary conditions given in equation (16) in the disk-slice $\Omega$. In Fig. 3, we also show the norms of each block of calculated coefficients of the approximation for four right-hand sides of the Poisson equation with $N=200$, that is, $(N+1)^{2}=40,401$ unknowns. The right-hand sides we choose here are given by

$$
f(x, y, z)=\left\|(x-(\epsilon+1 / \sqrt{3}), y-(\epsilon+1 / \sqrt{3}), z-(\epsilon+1 / \sqrt{3}))^{\top}\right\|
$$

for differing choices of $\epsilon$-this parameter serves to alter the distance from which we would have a singularity. In the plot, a 'block' is simply the group of coefficients corresponding to OPs of the same degree, and so the plot shows how the norms of these blocks decay as the degree of the expansion increases. Thus, the rate of decay in the coefficients is a proxy for the rate of convergence of the


FIG. 3. Top: the computed solution to $\Delta u=f$ with zero boundary conditions with $f(x, y, z)=-2 e^{x} y z(2+x)+w_{R}^{(1,0)}(z) e^{x}\left(y^{3}+\right.$ $z^{2} y-4 x y-2 y$ ). Bottom: the norms of each block of the computed solution of the Poisson equation with right-hand side function $f(x, y, z)=\left\|\mathbf{x}-(\epsilon+1 / \sqrt{3})(1,1,1)^{\top}\right\|$ for different $\epsilon$ values. This indicates spectral convergence.
computed solution: as typical of spectral methods, we expect the numerical scheme to converge at the same rate as the coefficients decay. We see that we achieve spectral convergence for these examples.

### 5.2 Inhomogeneous variable-coefficient Helmholtz

Find $u(x, y)$ given functions $v, f: \Omega \rightarrow \mathbb{R}$ such that

$$
\begin{cases}\Delta_{\mathrm{S}} u(x, y, z)+k^{2} v(x, y, z) u(x, y, z)=f(x, y, z) & \text { in } \Omega  \tag{17}\\ u(x, y, z)=0 & \text { on } \partial \Omega\end{cases}
$$

where $k \in \mathbb{R}$, noting the imposition of zero Dirichlet boundary conditions on $u$.


FIg. 4. Top: the computed solution to $\Delta u+k^{2} v u=f$ with zero boundary conditions with $f(x, y, z)=y e^{x}(z-\alpha), v(x, y, z)=$ $1-\left(3\left(x-x_{0}\right)^{2}+5\left(y-y_{0}\right)^{2}+2\left(z-z_{0}\right)^{2}\right)$ where $\left(x_{0}, z_{0}\right):=(0.7,0.2), y_{0}:=\sqrt{1-x_{0}^{2}-z_{0}^{2}}$ and $k=100$. Bottom: the norms of each block of the computed solution of the Helmholtz equation with the right-hand side function $f(x, y, z)=1$ and the same function $v(x, y, z)$, for various $k$ values. This indicates spectral convergence.

We can tackle the problem as follows. Denote the coefficient vector for expansion of $u$ in the $\mathbb{W}_{N}^{(1)}$ OP basis up to degree $N$ by $\mathbf{u}$, and the coefficient vector for expansion of $f$ in the $\tilde{\mathbb{Q}}_{N}^{(1)}$ OP basis up to degree $N$ by $\mathbf{f}$. Since $f$ is known, we can obtain the coefficients $\mathbf{f}$ using the quadrature rule in Section 4.3.

Define $X:=\left(J_{x}^{(0)}\right)^{\top}, Y:=\left(J_{y}^{(0)}\right)^{\top}, Z:=\left(J_{z}^{(0)}\right)^{\top}$. We can obtain the matrix operator for the variablecoefficient function $v(x, y, z)$ by using the Clenshaw algorithm with matrix inputs as the Jacobi matrices $X, Y, Z$, yielding an operator matrix of the same dimension as the input Jacobi matrices a la the procedure introduced in Olver et al. (2019). We can denote the resulting operator acting on coefficients in the $\widetilde{\mathbb{Q}}_{N}^{(0)}$


FIG. 5. Time in seconds to build and solve the system $\left[\Delta_{\mathrm{S}}+v(x, y, z)\right] u(x, y, z)=f(x, y, z)$, for a rotationally invariant $v(x, y, z)=v(z)$. This demonstrates that the approach is roughly of order $\mathcal{O}\left(N^{2}\right)$, where $N$ is the degree to which we approximate the solution (the number of unknowns is then $\left.(N+1)^{2}\right)$. Here, we used $f=-2 e^{x} y z(2+x)+(z-\alpha) e^{x}\left(y^{3}+z^{2} y-4 x y-2 y\right)$ and $v(x, y, z)=v(z)=\cos (z)$.
space by $v(X, Y, Z)$. In matrix-vector notation, our system hence becomes

$$
\left(\Delta_{W}^{(1)}+k^{2} T^{(0) \rightarrow(1)} V T_{W}^{(1) \rightarrow(0)}\right) \mathbf{u}=\mathbf{f}
$$

which can be solved to find $\mathbf{u}$. We can see the sparsity and structure of this matrix system in Fig. 1 with $v(x, y, z)=z x y^{2}$ as an example. In Fig. 4, we see the solution to the inhomogeneous variablecoefficient Helmholtz equation with zero boundary conditions given in equation (17) in the spherical cap $\Omega$, with $f(x, y, z)=y e^{x} w_{R}^{(1,0)}(z), v(x, y, z)=1-\left(3\left(x-x_{0}\right)^{2}+5\left(y-y_{0}\right)^{2}+2\left(z-z_{0}\right)^{2}\right)$ where $\left(x_{0}, z_{0}\right):=(0.7,0.2), y_{0}:=\sqrt{1-x_{0}^{2}-z_{0}^{2}}$ and $k=100$. In Fig. 4, we also show the norms of each block of calculated coefficients for the approximation of the solution to the inhomogeneous variablecoefficient Helmholtz equation with various $k$ values. Here, we use $N=200$, that is, $(N+1)^{2}=40,401$ unknowns. Once again, the rate of decay in the coefficients is a proxy for the rate of convergence of the computed solution, and we see that we achieve spectral convergence.

In Fig. 5, we plot the time taken ${ }^{2}$ to construct the operator for $\Delta_{\mathrm{S}}+v(x, y, z)$, with a rotationally invariant $v(x, y, z)=v(z)=\cos z$ and solve a zero boundary condition Helmholtz problem. The plot demonstrates that as we increase the degree of approximation $N$, we achieve a complexity of an optimal $\mathcal{O}\left(N^{2}\right)$.

What about other boundary conditions? One simple extension is the case where the value on the boundary takes that of a function depending only on $x$ and $y$, i.e., $c=c(x, y)$. In this case, the problem

$$
\begin{cases}\Delta_{\mathrm{S}} u(x, y, z)+k^{2} v(x, y, z) u(x, y, z)=f(x, y, z) & \text { in } \Omega \\ u(x, y, z)=c(x, y) & \text { on } \partial \Omega\end{cases}
$$

[^2]is equivalent to letting $u(x, y, z)=\tilde{u}(x, y, z)+c(x, y)$ and solving
\[

\left\{$$
\begin{array}{l}
\Delta_{\mathrm{S}} \tilde{u}(x, y, z)+k^{2} v(x, y, z) \tilde{u}(x, y, z)=f(x, y, z)-k^{2} v(x, y, z) c(x, y)-\Delta_{\mathrm{S}} c(x, y) \quad \text { in } \Omega \\
\tilde{u}(x, y, z)=0 \text { on } \partial \Omega
\end{array}
$$\right.
\]

for $\tilde{u}$. This new problem is then a zero boundary condition Helmholtz problem with right-hand side

$$
g(x, y, z):=f(x, y, z)-k^{2} v(x, y, z) c(x, y)-\Delta_{\mathrm{S}} c(x, y)
$$

for $(x, y, z) \in \Omega$. Notice that the spherical Laplacian applied to $c(x, y)$, expanded in the $\widetilde{\mathbb{Q}}_{N}^{(1)}$ basis with coefficients vector $\boldsymbol{c}=\left(c_{n, k, i}\right)$, is just

$$
\Delta_{\mathrm{S}} c(x, y)=\frac{1}{\rho(z)^{2}} \sum_{n=0}^{N} \sum_{i=0}^{1} c_{n, n, i} \frac{\partial^{2}}{\partial \theta^{2}} Y_{n, i}(\theta)=-\frac{1}{\rho(z)^{2}} \sum_{n=0}^{N} \sum_{i=0}^{1} n^{2} c_{n, n, i} Y_{n, i}(\theta)
$$

since the coefficients $\left\{c_{n, k, i}\right\}$ for such a function are zero for $k<n$ due to the dependence on $x$ and $y$ only, which are precisely the Fourier coefficients of $c(\cos \theta, \sin \theta)$. Thus, since the function $c(x, y)$ is known, it is simple to evaluate $\frac{\partial^{2}}{\partial \theta^{2}} c(x, y)$ and hence one can obtain the coefficients for the expansion of $g(x, y, z)$ in the $\tilde{\mathbb{Q}}_{N}^{(1)}$ basis in the usual manner.

### 5.3 Biharmonic equation

Our last example is the biharmonic equation: find $u(x, y, z)$ given a function $f(x, y, z)$ such that

$$
\begin{cases}\Delta_{\mathrm{S}}^{2} u(x, y, z)=f(x, y, z) & \text { in } \Omega  \tag{18}\\ u(x, y, z)=0, \quad \frac{\partial u}{\partial n}(x, y, z)=\nabla_{S} u(x, y, z) \cdot \hat{\boldsymbol{n}}(x, y, z)=0 & \text { on } \partial \Omega\end{cases}
$$

where $\Delta_{\mathrm{S}}^{2}$ is the Biharmonic operator, noting the imposition of zero Dirichlet and Neumann boundary conditions on $u$. For clarity, we reiterate that the unit normal vector in this sense is simply $\hat{\boldsymbol{n}}(x, y, z)=$ $\hat{\boldsymbol{n}}(\boldsymbol{x}):=\frac{\boldsymbol{x}}{\|\boldsymbol{x}\|}=\boldsymbol{x}$ (see Section 3.1). In Fig. 6, we see the solution to the Biharmonic equation (18) in the spherical cap $\Omega$. In Fig. 6, we also show the norms of each block of calculated coefficients of the approximation for four more complex right-hand sides of the biharmonic equation with $N=200$, that is, $(N+1)^{2}=40,401$ unknowns. Once again, the rate of decay in the coefficients is a proxy for the rate of convergence of the computed solution, and we see that we achieve exponential convergence for these more complex functions.

### 5.4 Other boundary conditions

Other boundary conditions beyond those already presented here are beyond the scope of this paper. However, the approach established for the case of a triangle domain in two dimensions shows how Neumann conditions can be considered as a system of equations in partial derivatives involving Dirchlet conditions (Olver et al., 2019). This approach may also work here for the spherical cap. Ideally, one would properly tackle the tangent space (which again is beyond the scope of this paper), which would lower the degrees of freedom.


Fig. 6. Left: the computed solution to $\Delta^{2} u=f$ with zero Dirichlet and Neumann boundary conditions with $f(x, y, z)=(1+$ $\left.\operatorname{erf}\left(5\left(1-10\left((x-0.5)^{2}+y^{2}\right)\right)\right)\right) \rho(z)^{2}$. Right: the norms of each block of the computed solution of the biharmonic equation with the right-hand side function $f(x, y, z)=\exp \left(-\epsilon\left(\left(x-x_{0}\right)^{2}+\left(y-y_{0}\right)^{2}+\left(z-z_{0}\right)^{2}\right)\right)$ where $\left(x_{0}, z_{0}\right):=(0.7,0.2), y_{0}:=\sqrt{1-x_{0}^{2}-z_{0}^{2}}$, for various $\epsilon$ values. This demonstrates algebraic convergence.

## 6. Conclusions

We have shown that trivariate OPs can lead to sparse discretizations of general linear PDEs on spherical cap domains, with Dirichlet boundary conditions on the $z=\alpha \in(0,1)$ boundary. We have provided a detailed practical framework for the application of the methods described for quadratic surfaces of revolution (Olver \& Xu, 2020), by utilizing the non-classical one-dimensional OPs on the interval $[\alpha, 1]$ with the weight $(z-\alpha)^{a}\left(1-z^{2}\right)^{b / 2}$ defined for the disk-slice case (Snowball \& Olver, 2020). Generalization to spherical bands ( $\alpha \leq z \leq \beta$ ) is straightforward. This work thus forms a building
block in developing an $h p$-finite element method to solve PDEs on the sphere by using spherical band and spherical cap shaped elements.

This work also serves as a stepping stone to constructing similar methods to solve partial differential equations on other three-dimensional sub-domains of the sphere such as spherical triangles. It is clear from the construction in this paper that discretizations of spherical gradients and Laplacian's are sparse on other suitable sub-components of the sphere. The resulting sparsity in high-polynomial degree discretizations presents an attractive alternative to methods based on bijective mappings (e.g., Bonev et al., 2018; Boyd, 2005; Shipton et al., 2018). Constructing these sparse spectral methods for surface PDEs on spherical triangles is future work and has applications in weather prediction (Staniforth \& Thuburn, 2012), though it is not yet clear how to directly construct, or perhaps compute, the necessary OPs.

The next stage is to develop an orthogonal basis for the tangent space of the spherical cap (or band) and obtain sparse differential operators for gradient, divergence, etc. On the complete sphere, the vector spherical harmonics that form the orthogonal basis are simply the gradients and perpendicular gradients of the scalar spherical harmonics (Barrera et al., 1985), which has been used effectively for solving PDEs on the sphere (Lecoanet et al., 2019; Vasil et al., 2019); however, we do not have that luxury for the spherical cap or band, and hence the construction of a basis will not be as straightforward.

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[^1]:    ${ }^{1}$ https://github.com/snowball13/OrthogonalPolynomialFamilies.j1

[^2]:    ${ }^{2}$ measured using the '@belapsed’ macro from the BenchmarkTools.jl package (Chen \& Revels, 2016) in Julia.

